



CMB instrumentation and detectors

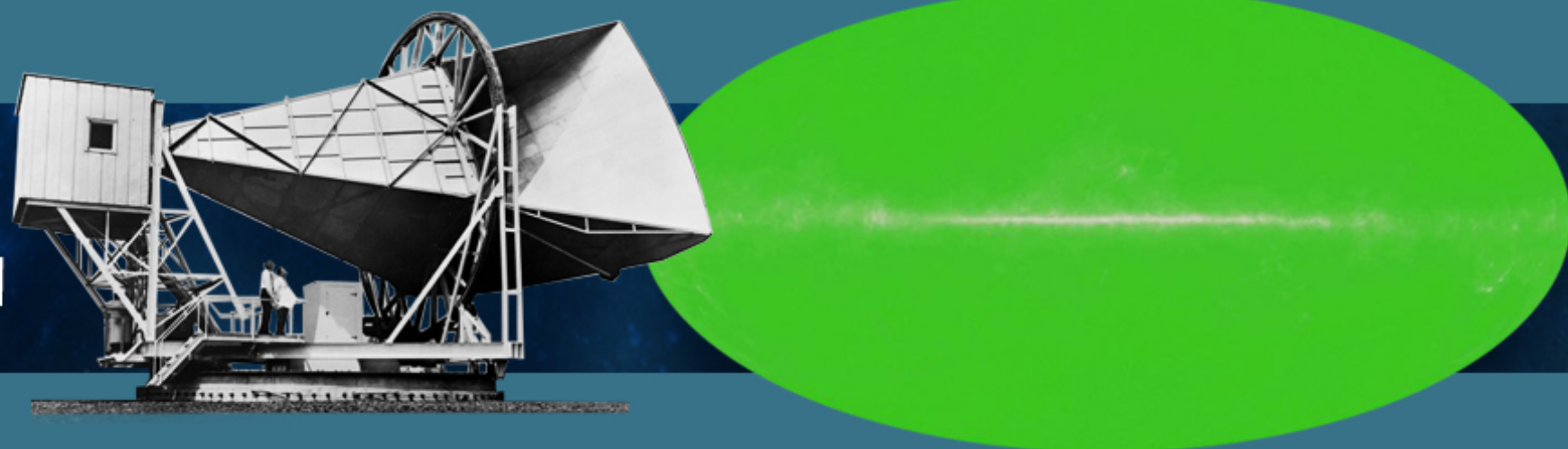
Thuong Duc HOANG

University of Science and Technology of Hanoi (USTH), Vietnam

June 8 2021

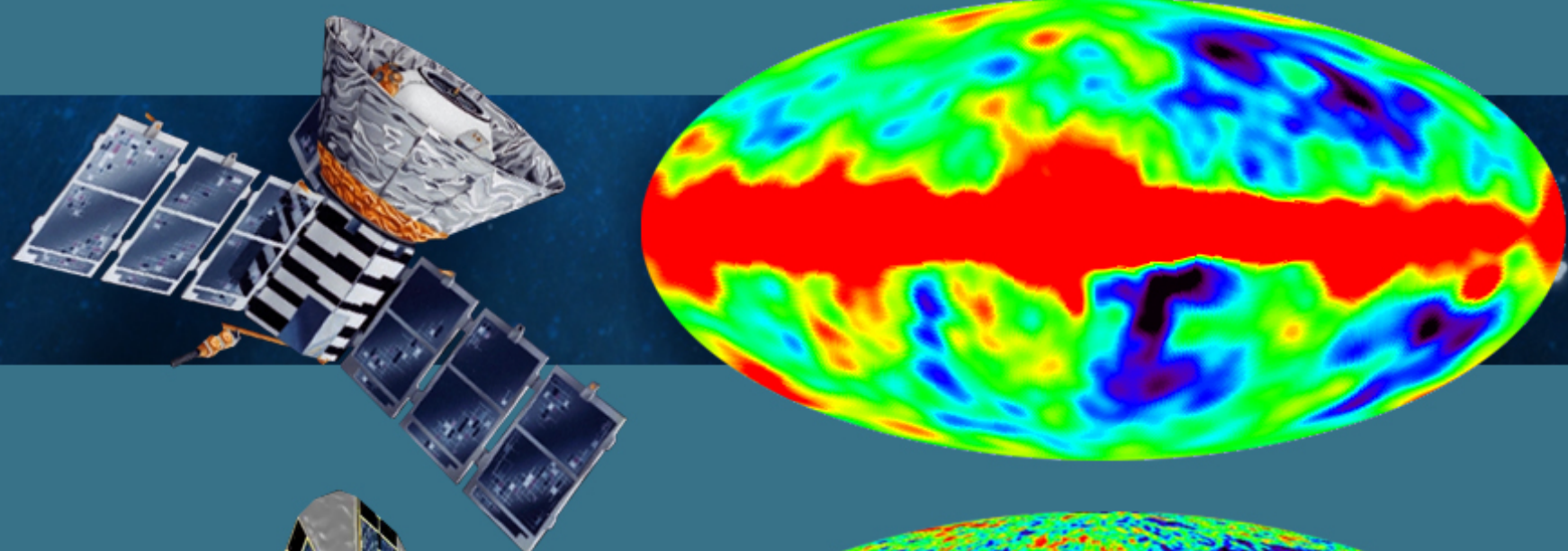
Horn antenna

1962
PENZIAS & WILSON



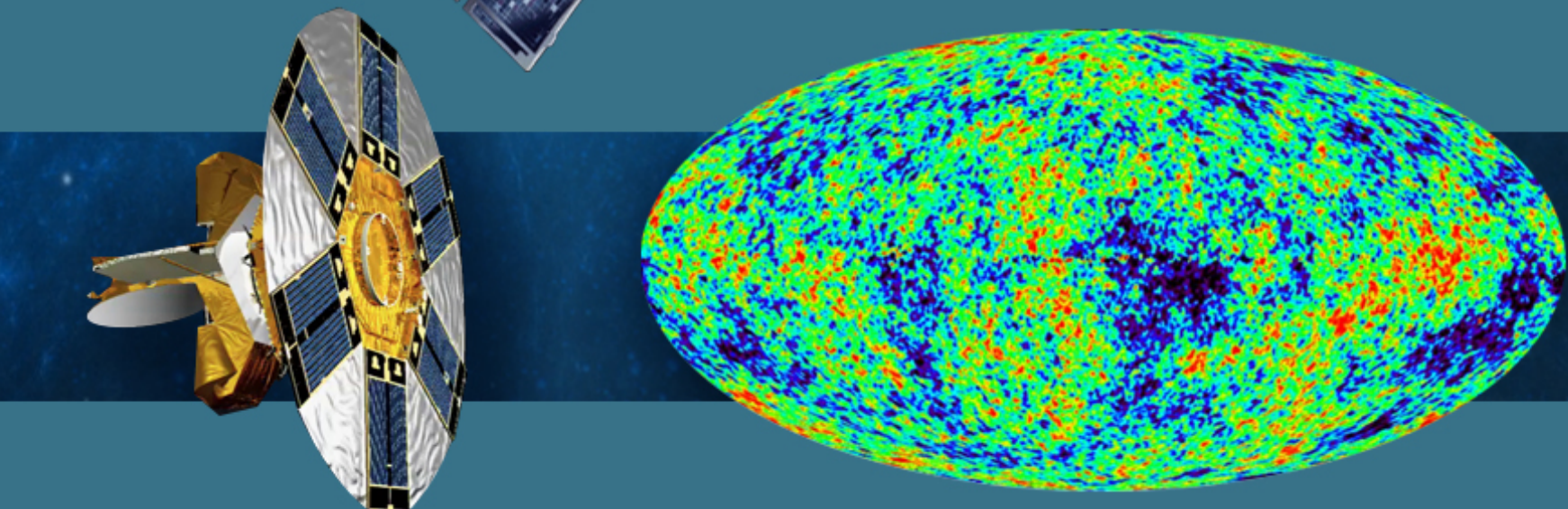
Microwaves
Radiometer

1989-1993
COBE



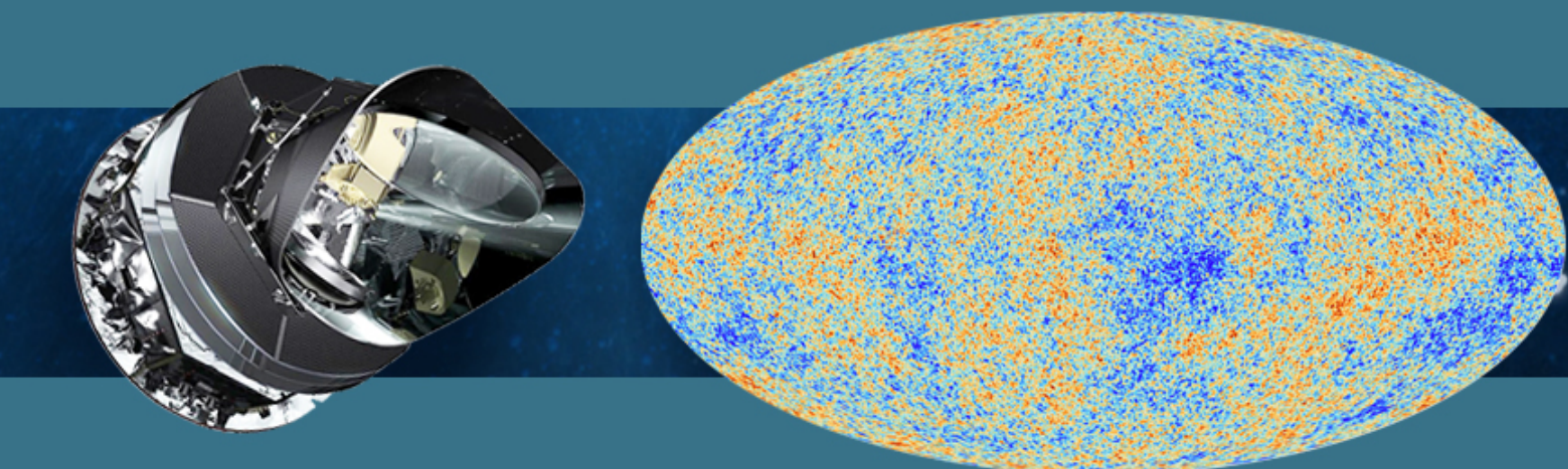
Microwaves
Radiometer

2001-2010
WMAP



Antenna
+ bolometers
(54 detectors)

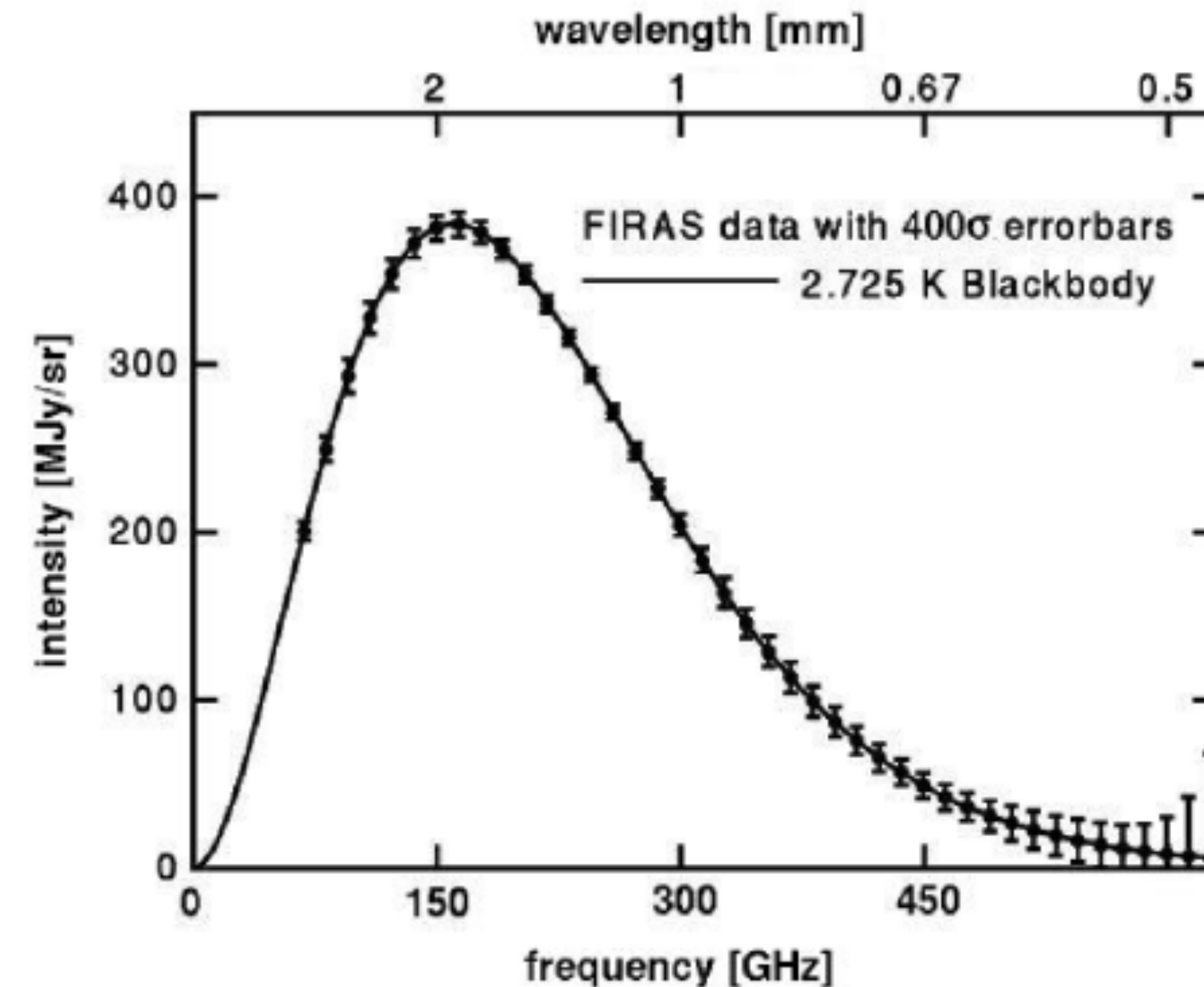
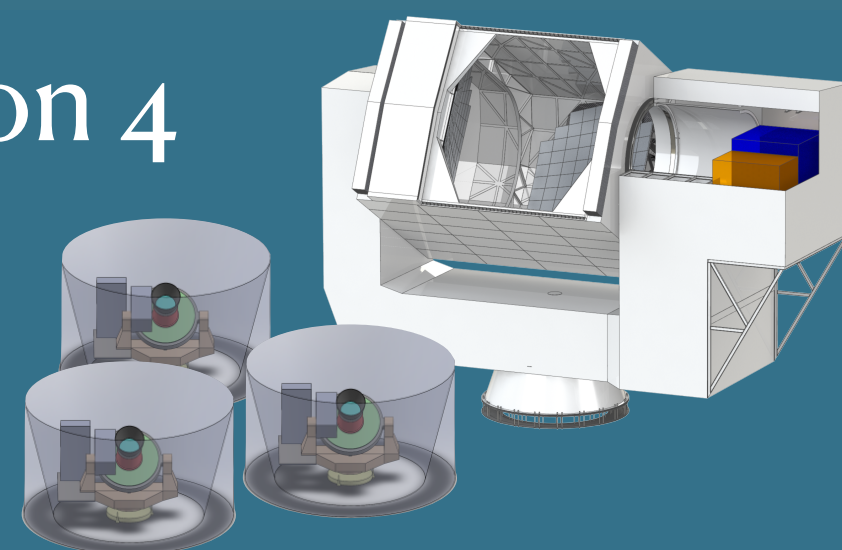
2009-2013
PLANK



Thousand
Superconducting
Transition Edge
Sensor (TES)



Generation 4



•The energy of CMB photons

$$E_{\text{CMB}} = \sim 1 \text{ meV}$$

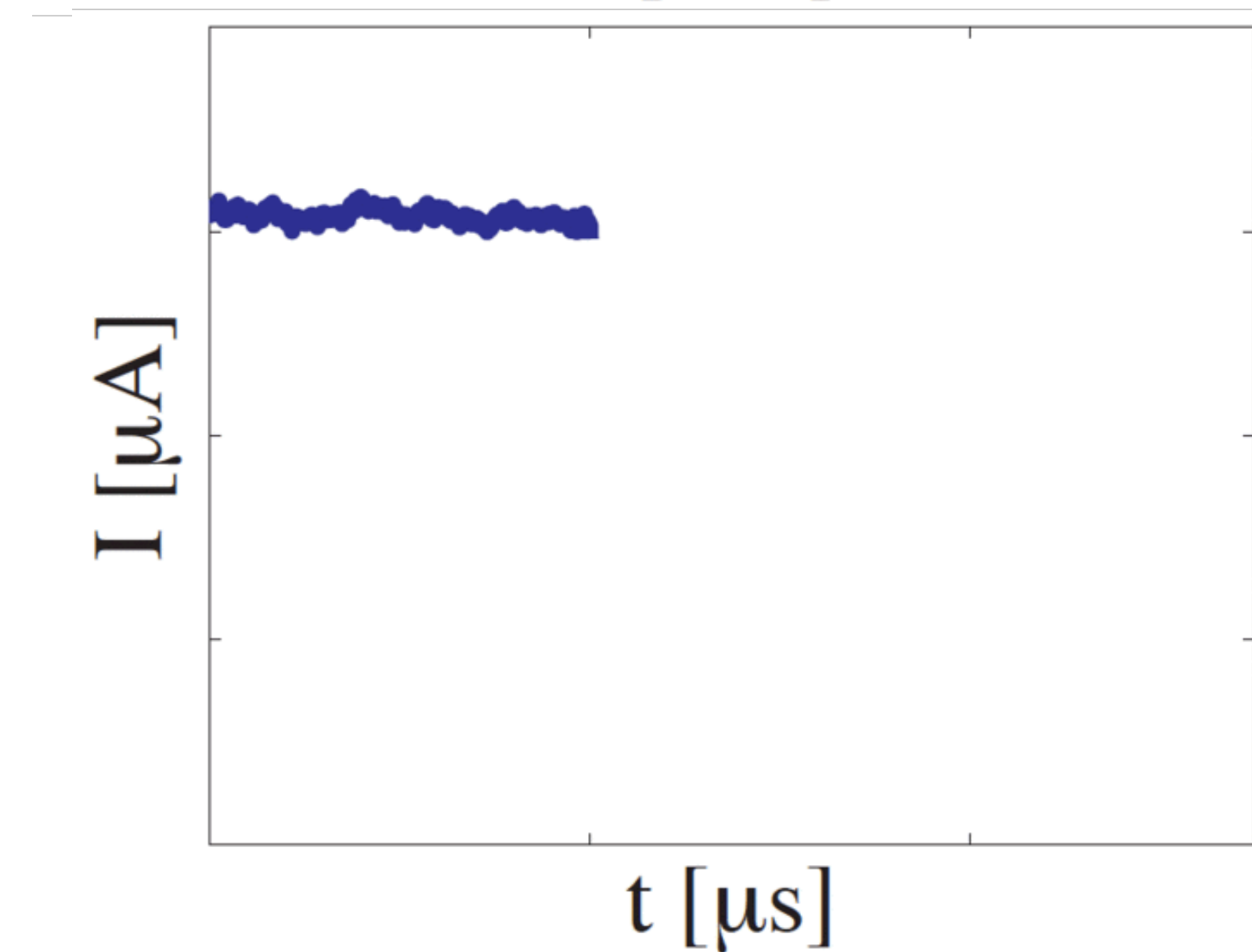
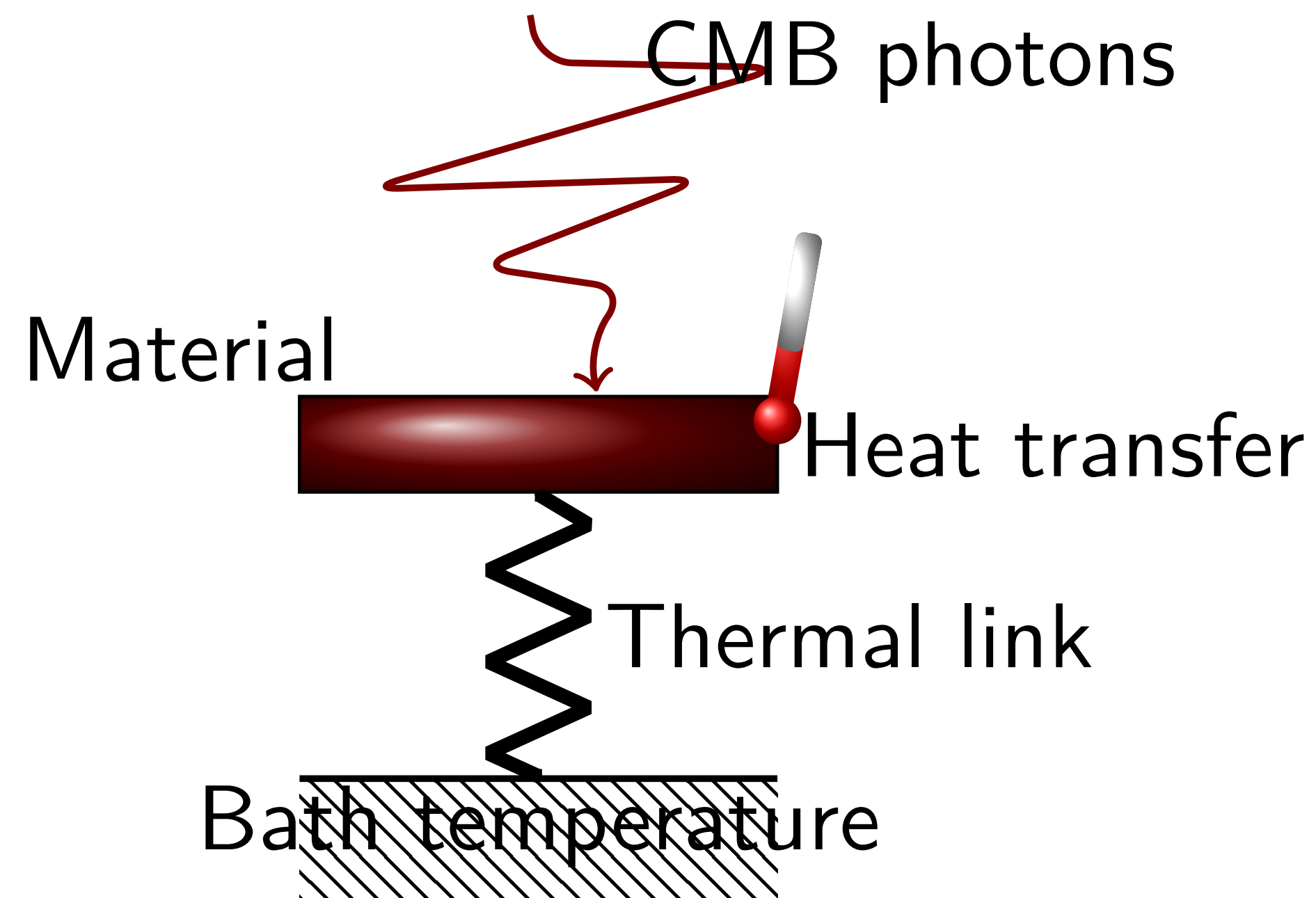
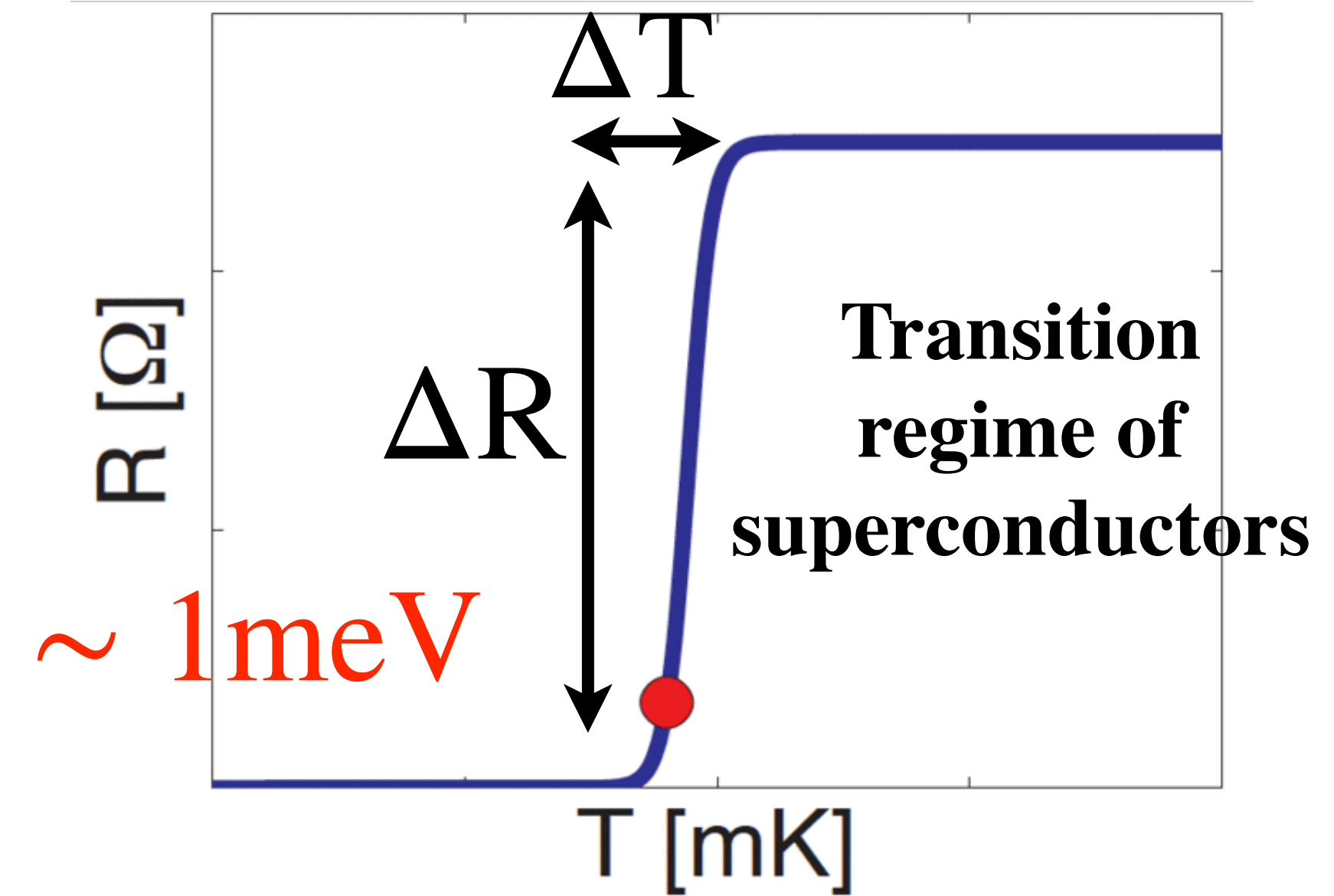
•CMB is a black body, the
peak frequency

$$\nu_{\text{max}} \sim 150 \text{ GHz or}$$

$$\lambda_{\text{max}} \sim 2 \text{ mm.}$$

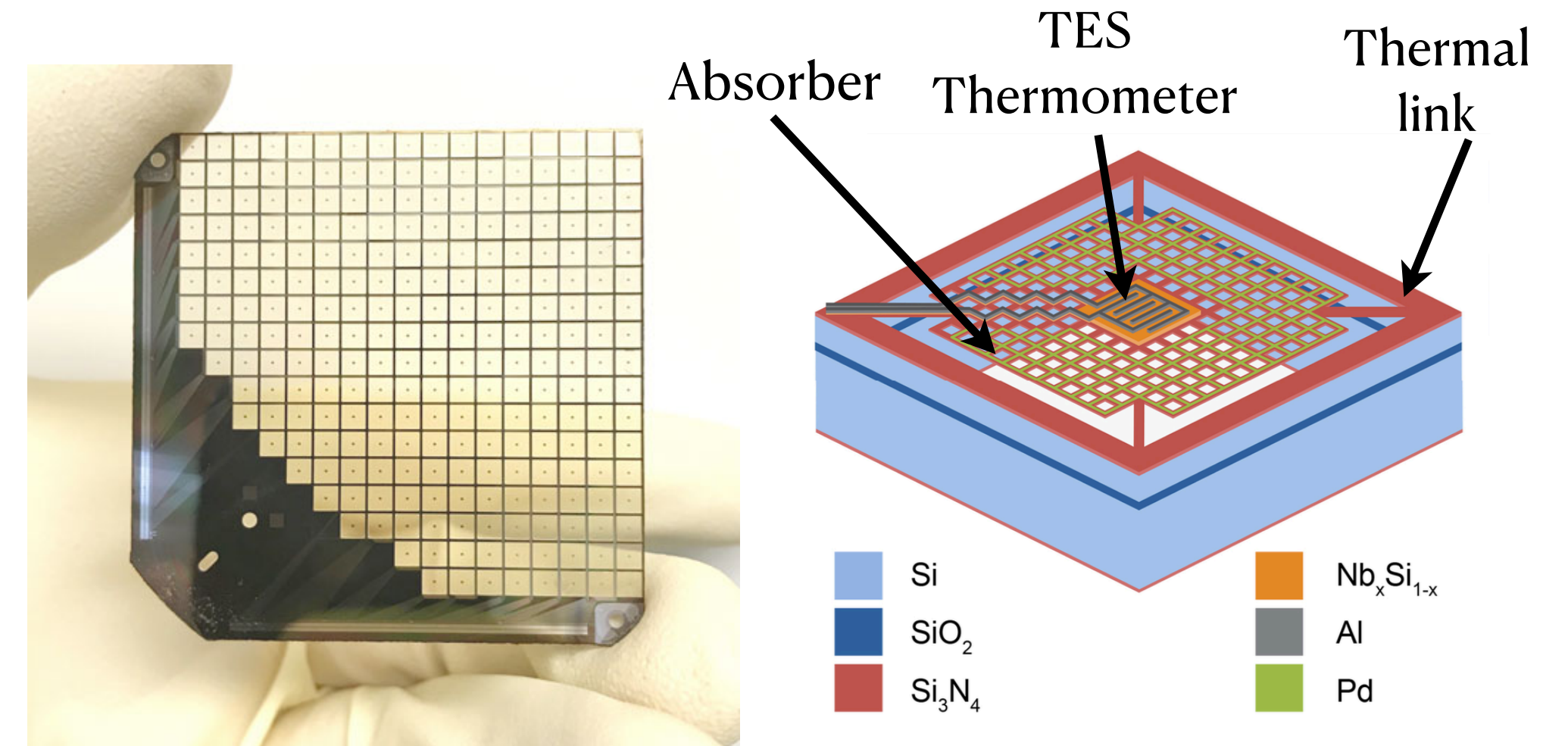
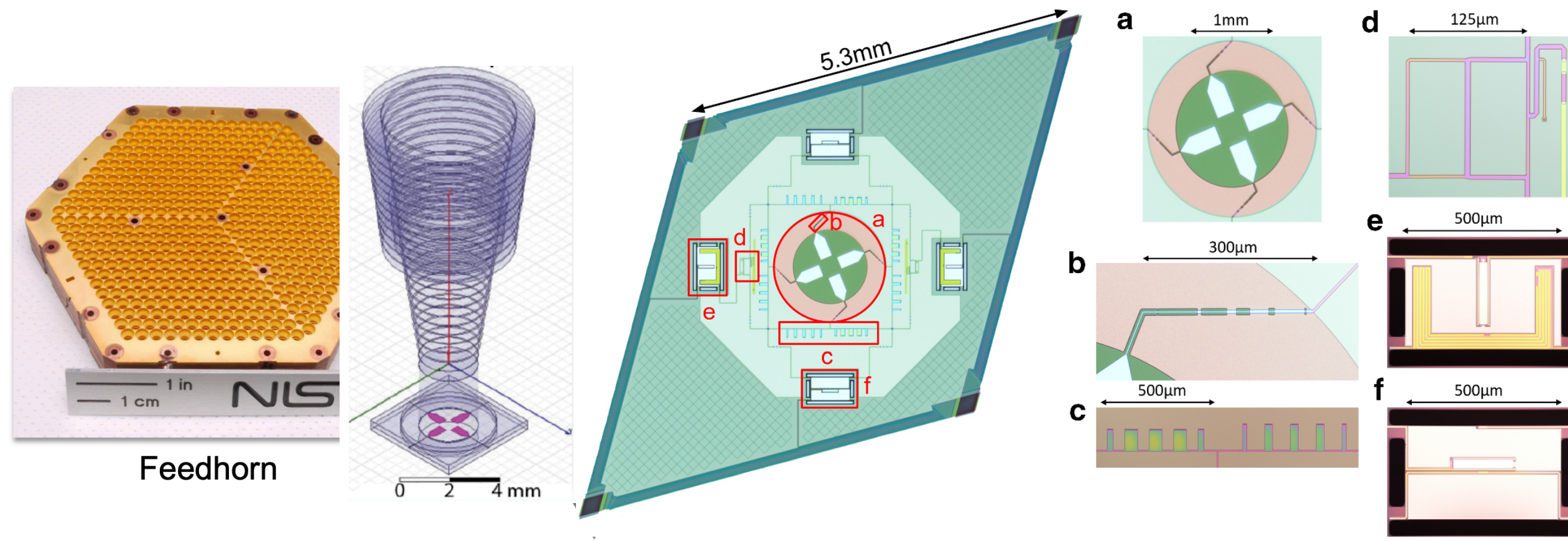
Superconducting Transition Regime

Temperature measurement!



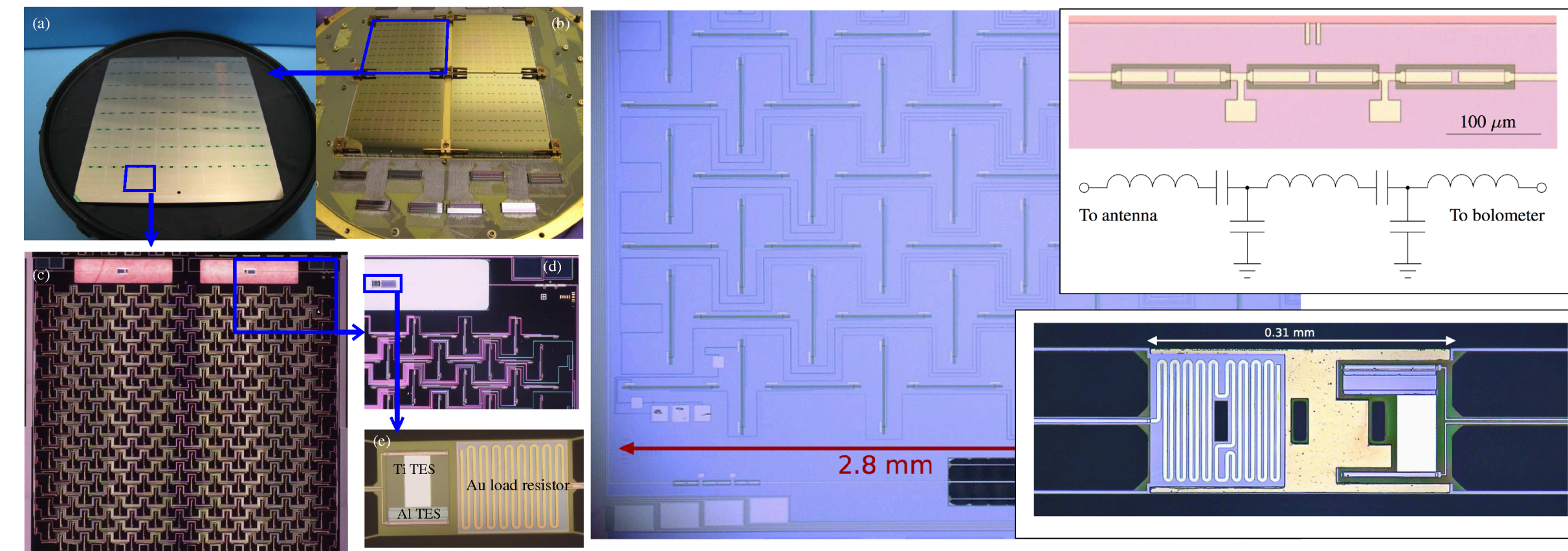
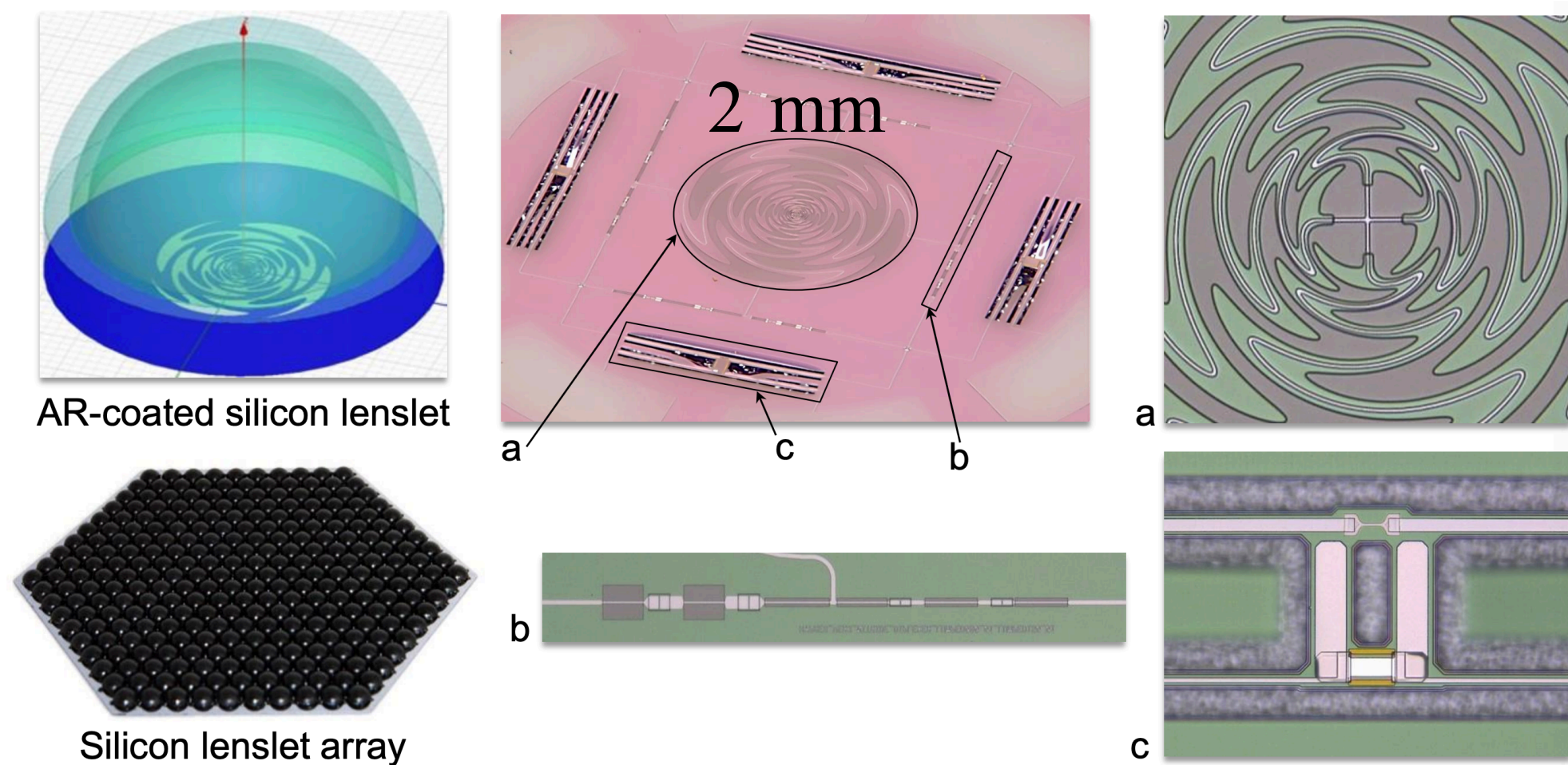
- **Transition Edge Sensor (TES):** A thermometer made from a superconducting film operated near its transition temperature.

TES Array Technology (mm scale)



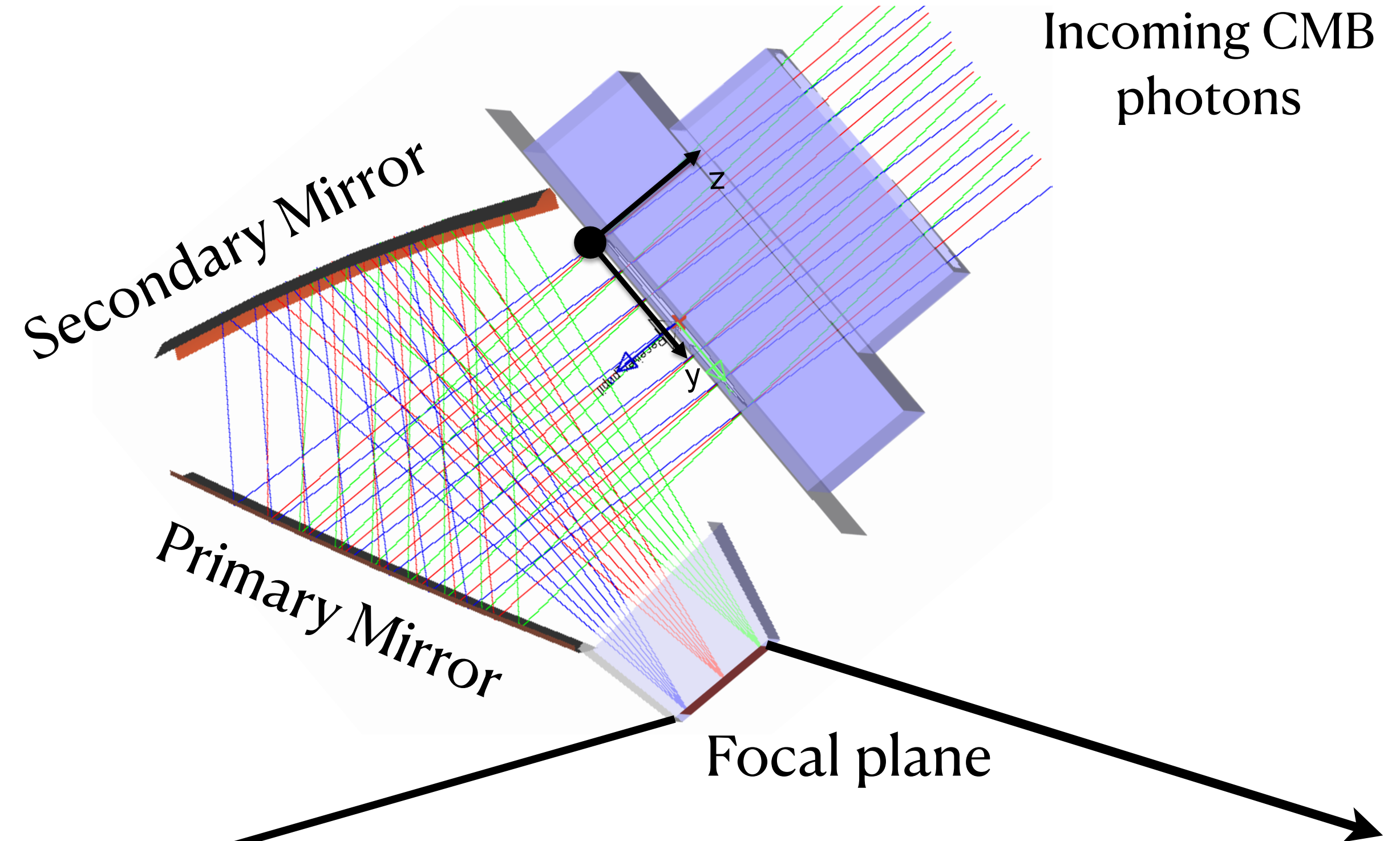
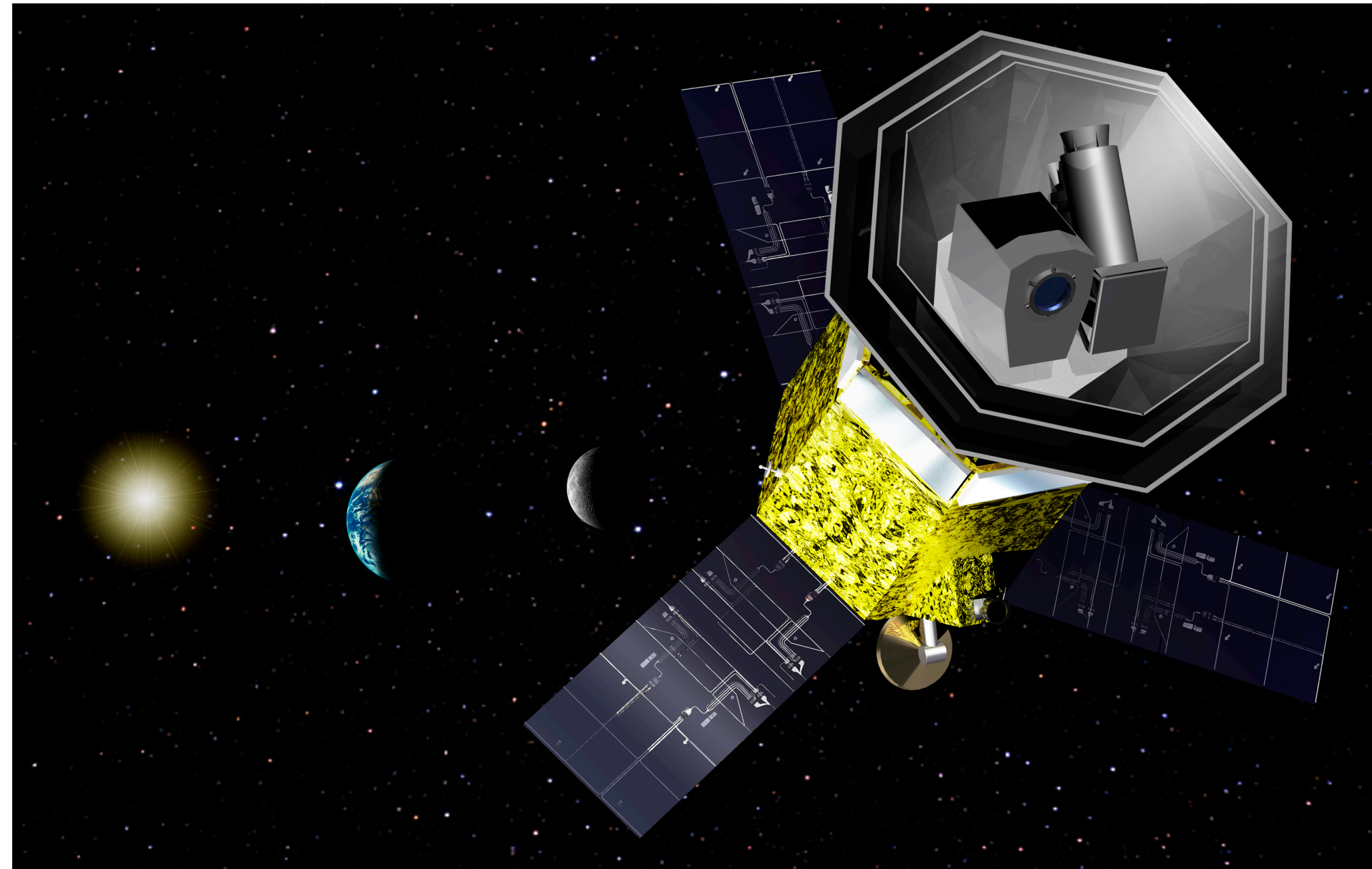
An array of 256 pixels, NbSi superconducting material (450 mK)
APC-France: no antenna technology: QUBIC

NIST: feed-horn technology: ACT, SPT, LiteBIRD (High frequency)

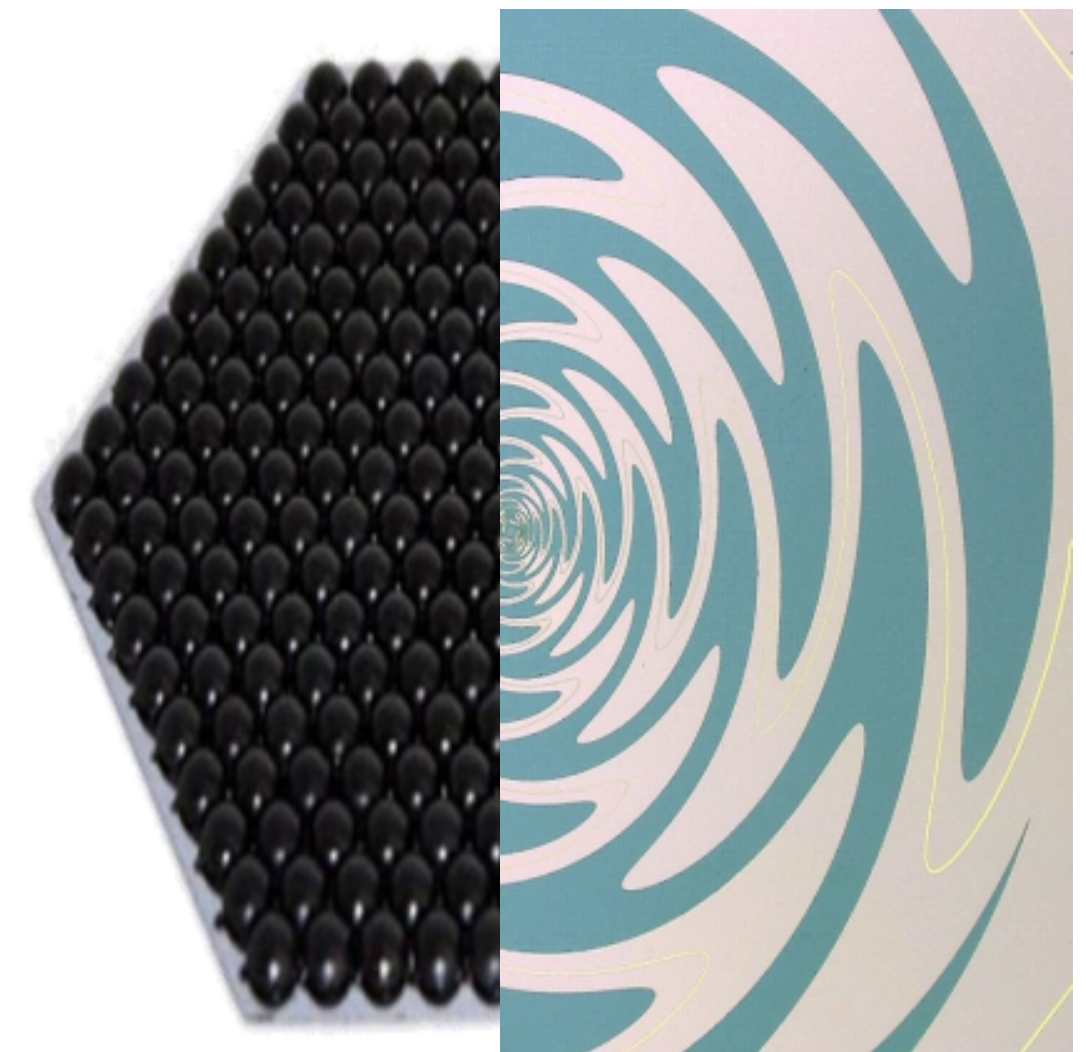
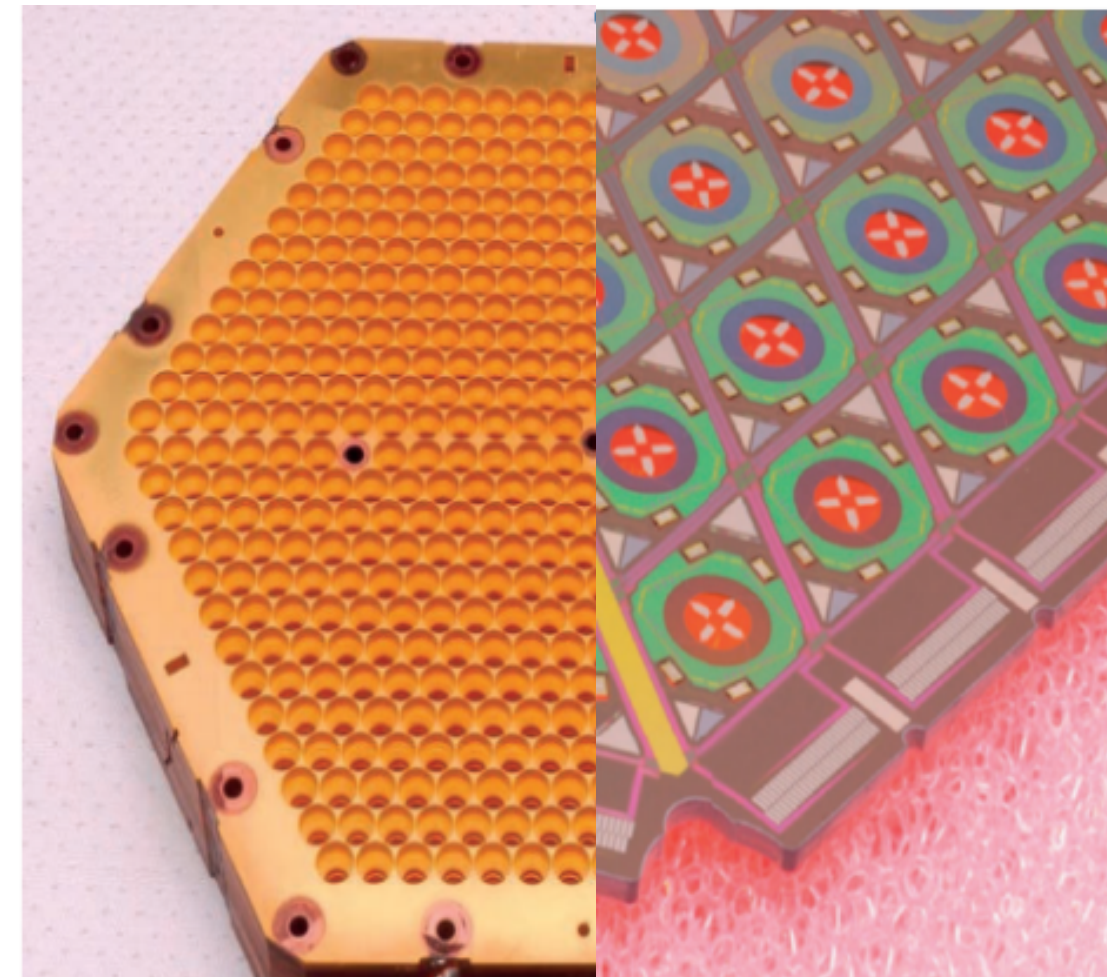


Berkeley: Planar sinuous antenna coupled TES: ACT, POLARBEAR, LiteBIRD (low-mid frequency)

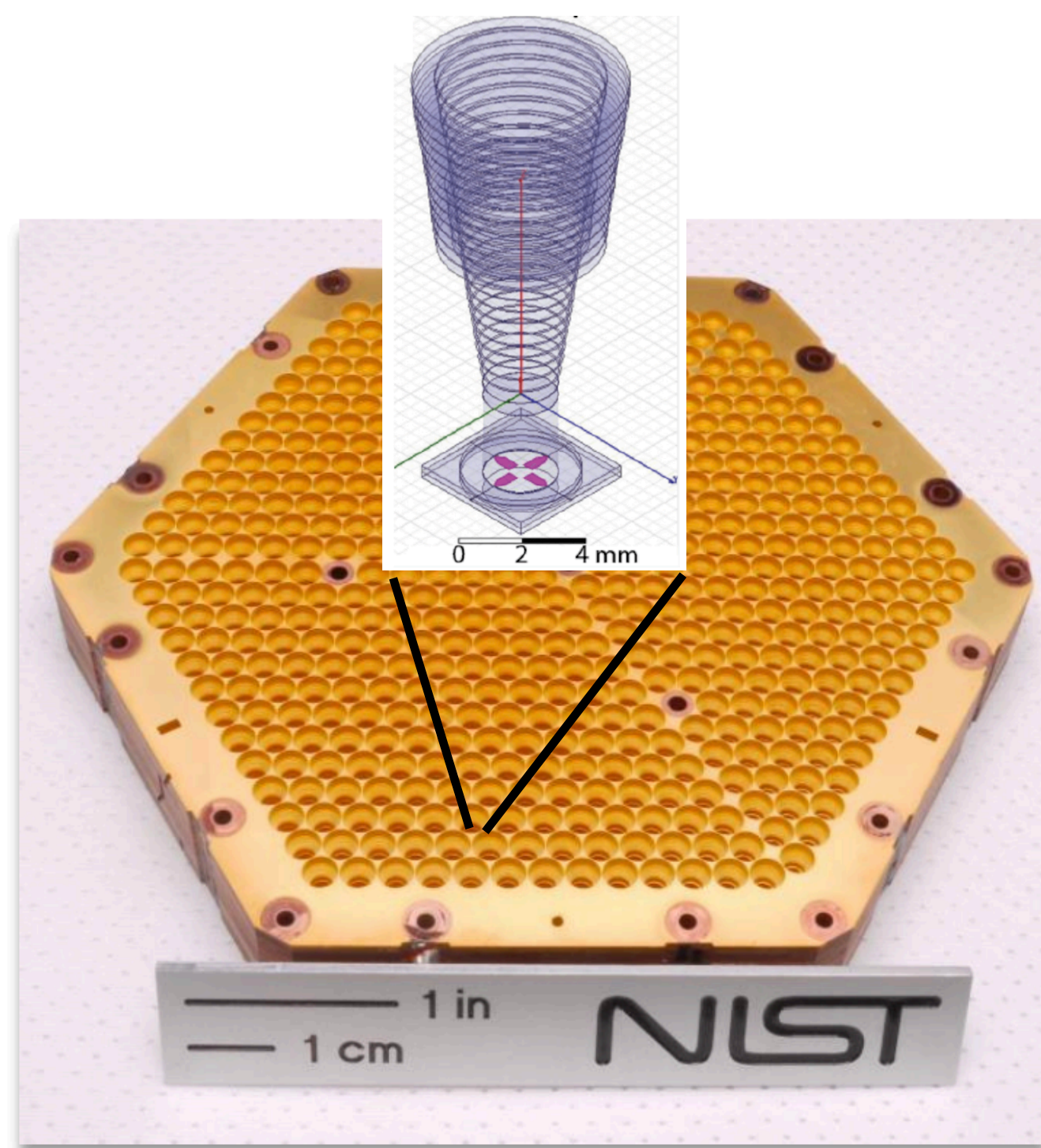
Caltech: Planar antenna coupled TES: BICEP/KECK



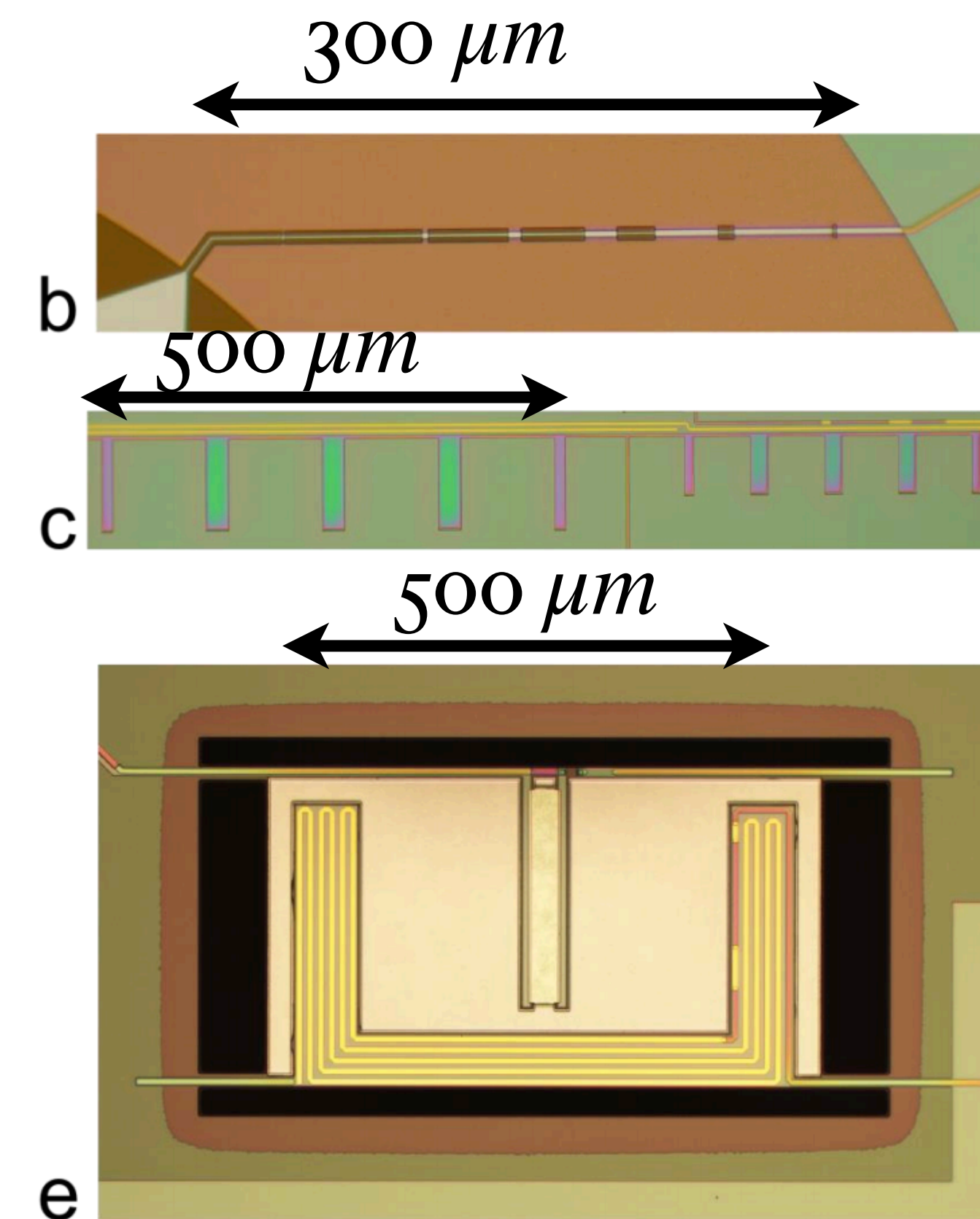
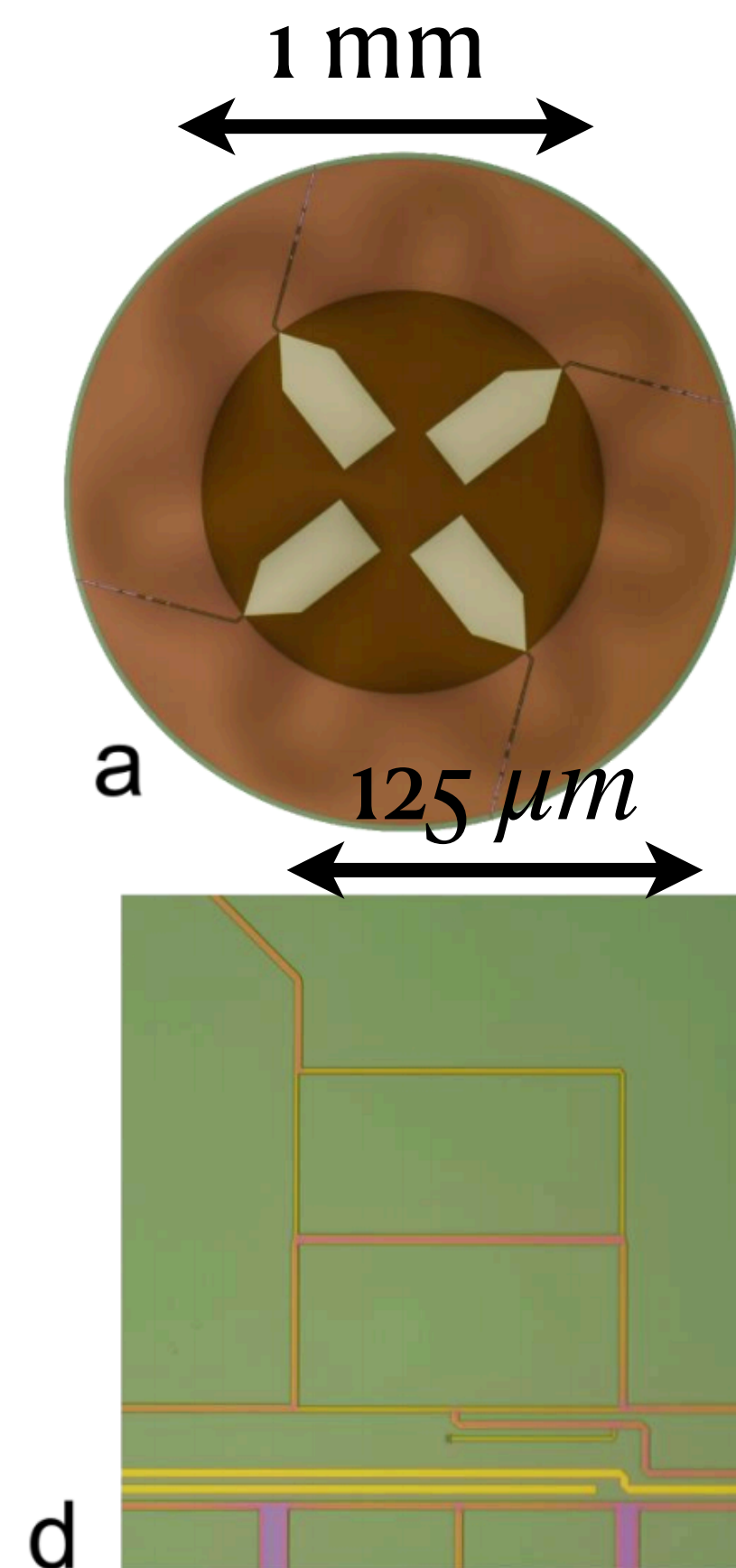
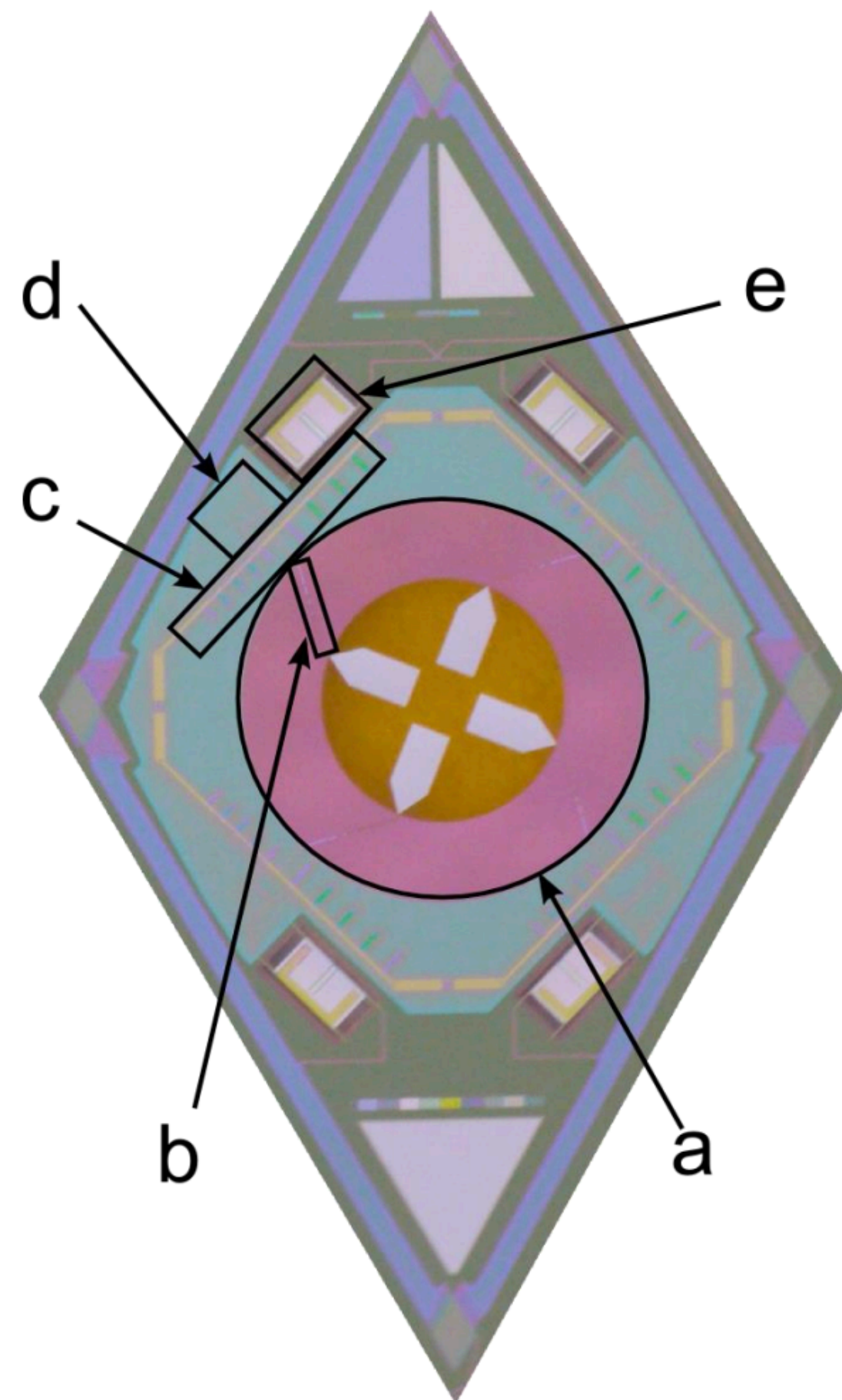
- **Two detector architectures: Feed horn and sinuous antenna**
- **Thousand TES detectors in the focal plane**



Horn-couple TES



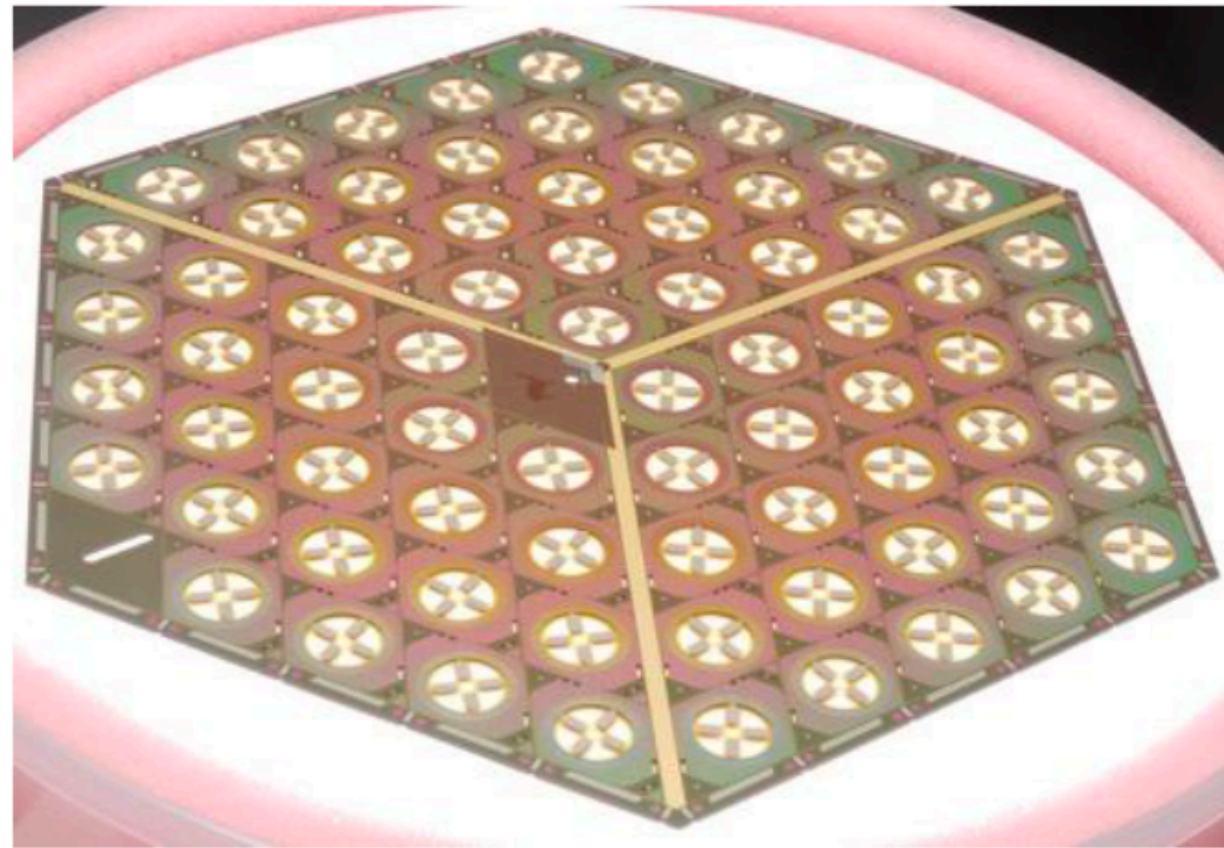
Feedhorn



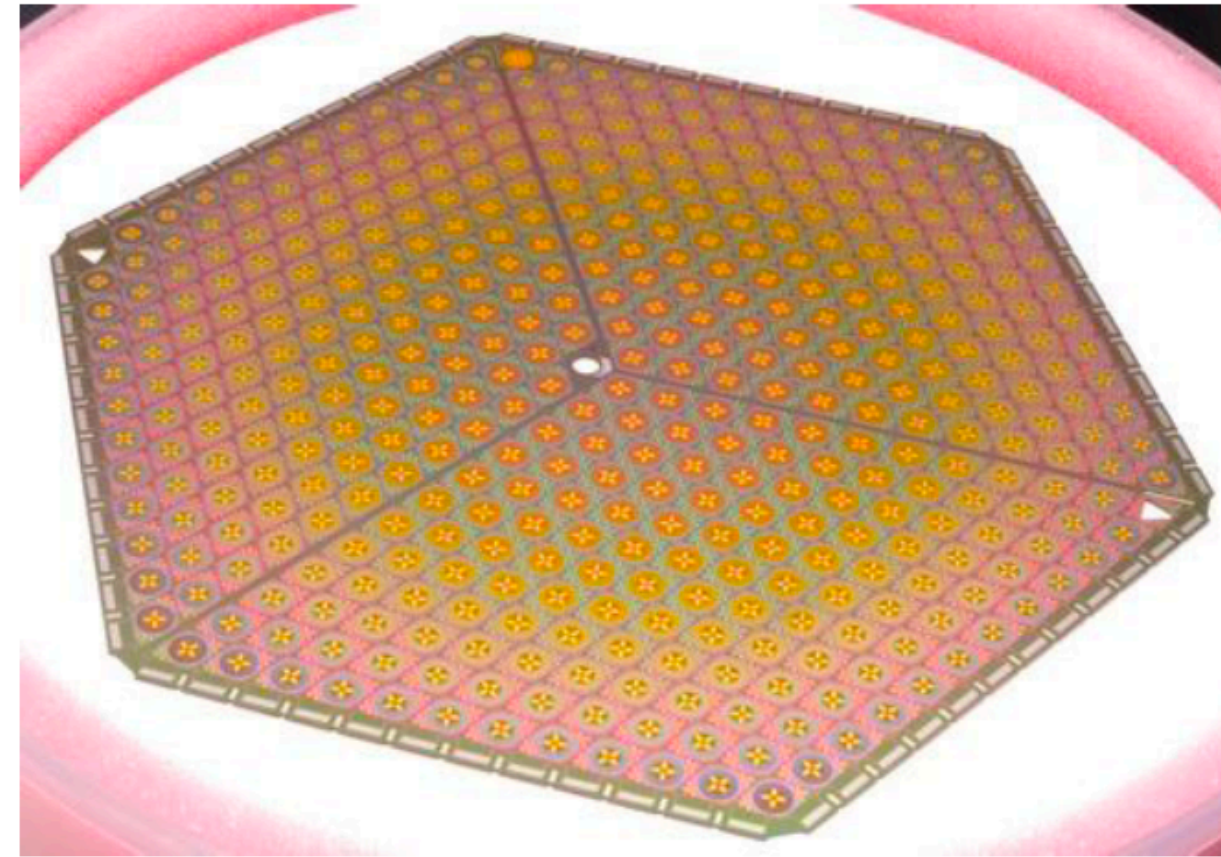
- a) A planar polarization antenna.
- b) Transmission lines.
- c) Bandpass filters for multifrequencies per chip.
- d) Hybrid tees which propagate the desirable mode.
- e) TES: AlMn (160 mK).

Horn-couple detectors

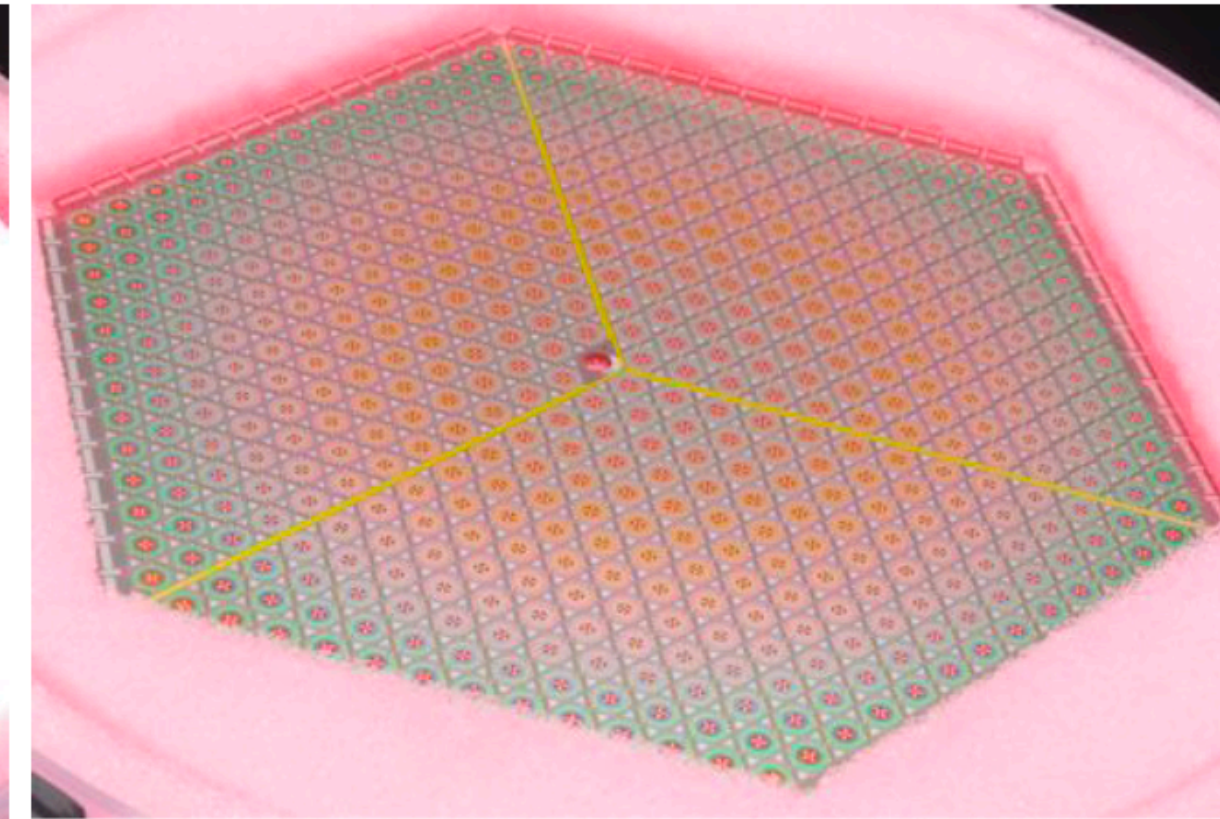
27/39GHz



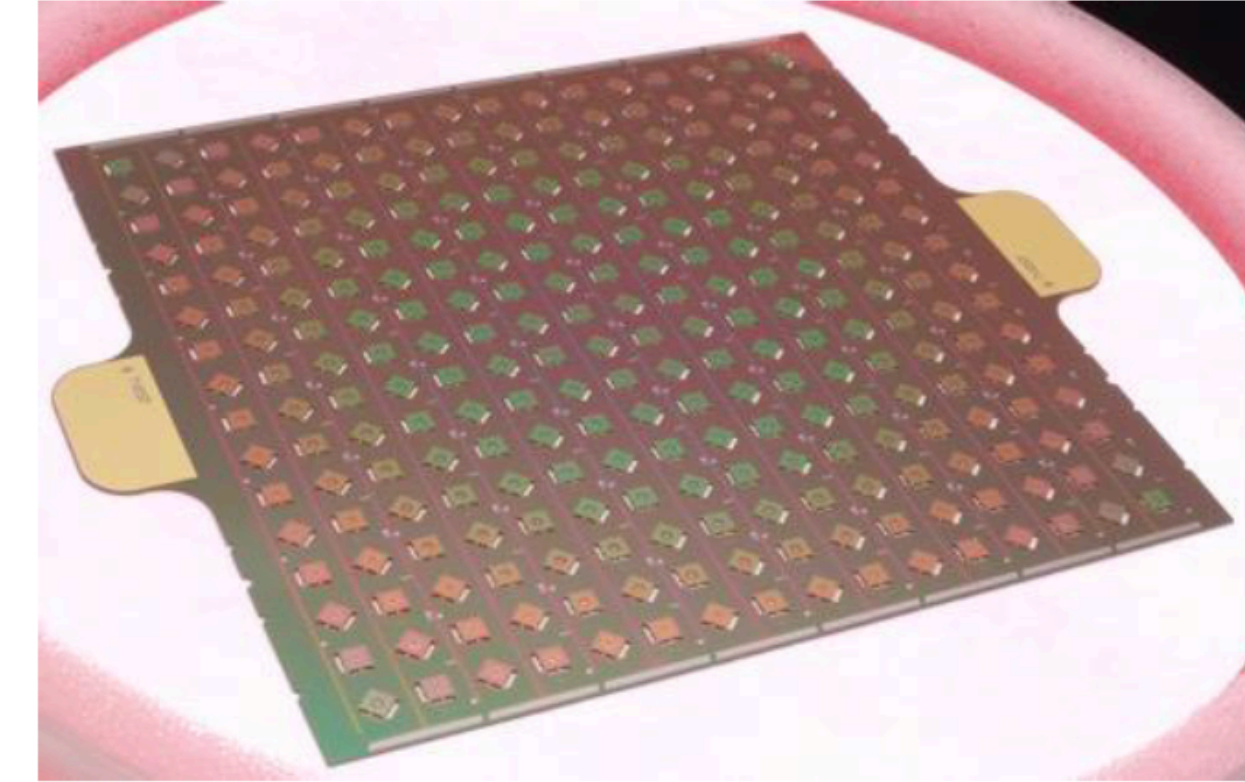
90/150GHz



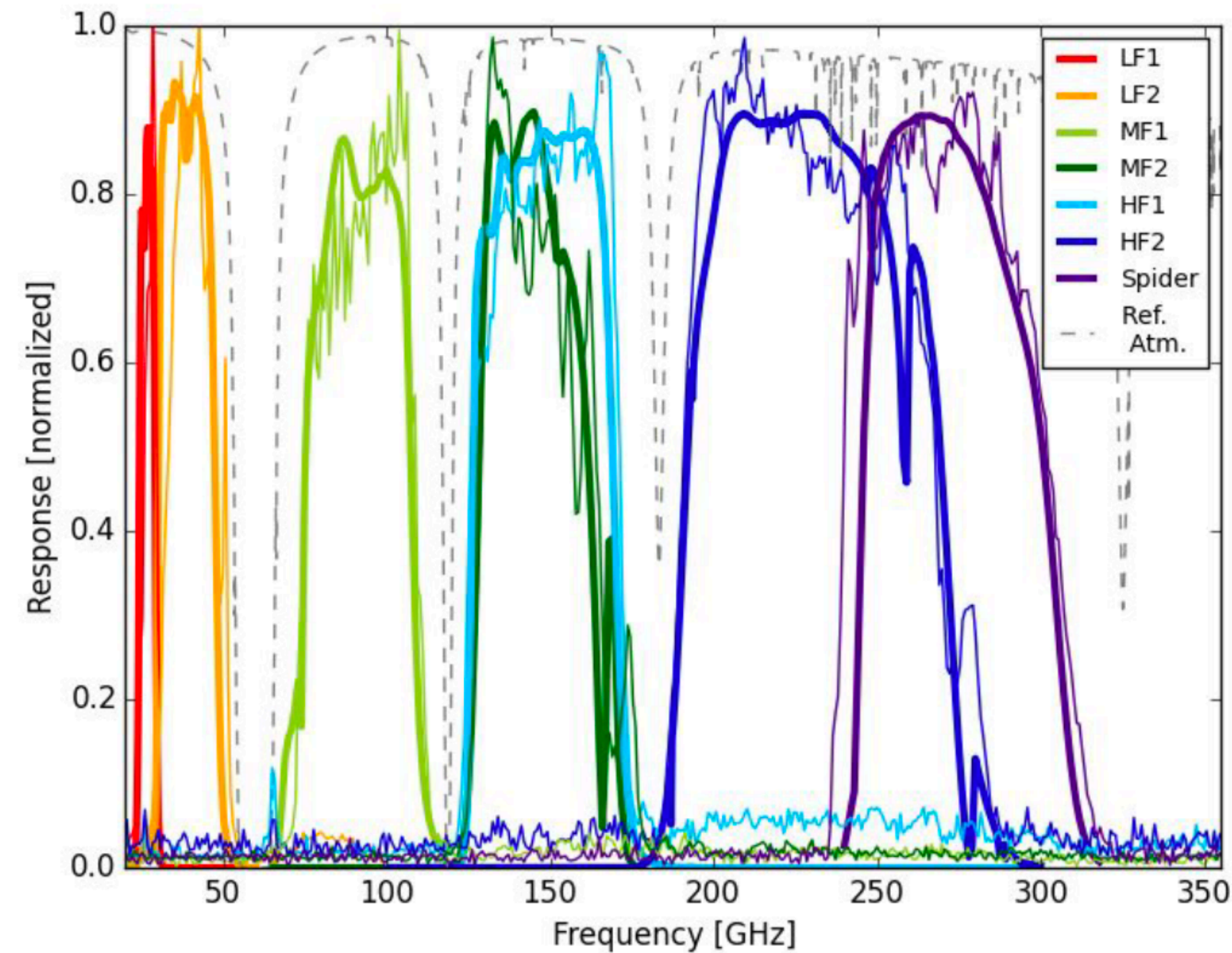
150/220GHz



280 GHz

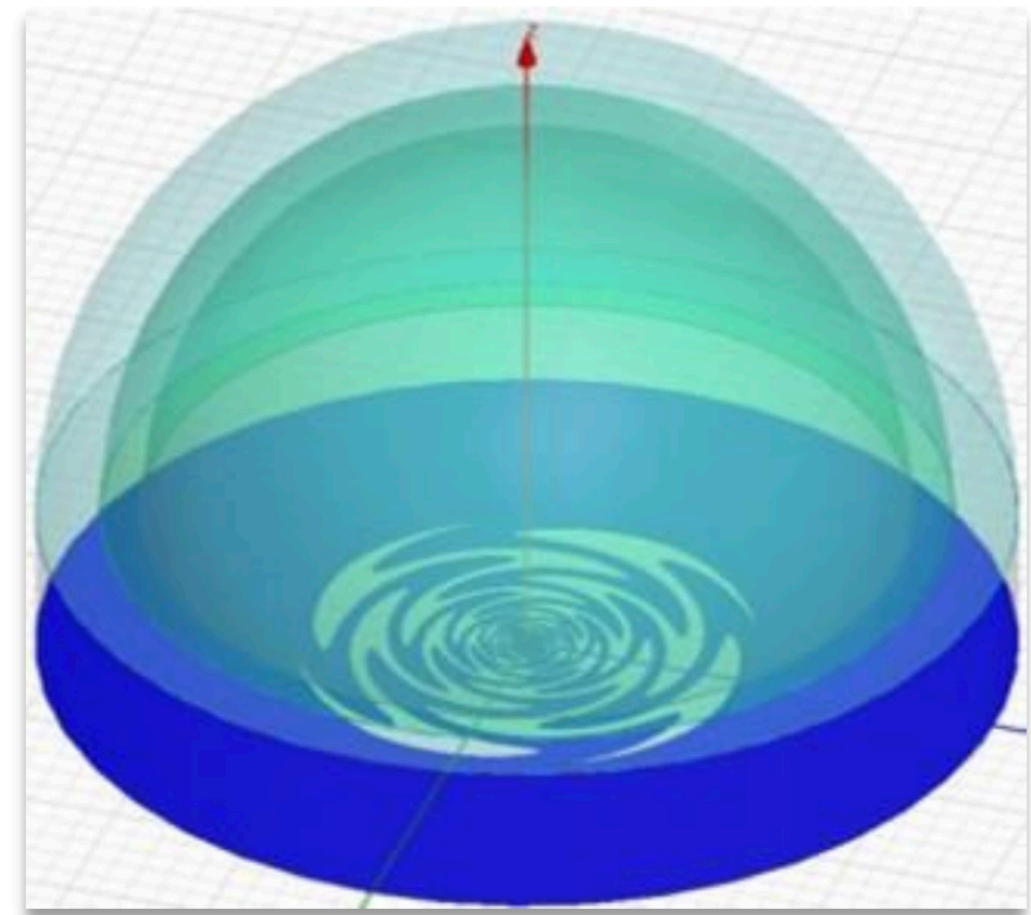


- A single wafer ~ 15 cm
- ~2000 TES
- Range of frequencies

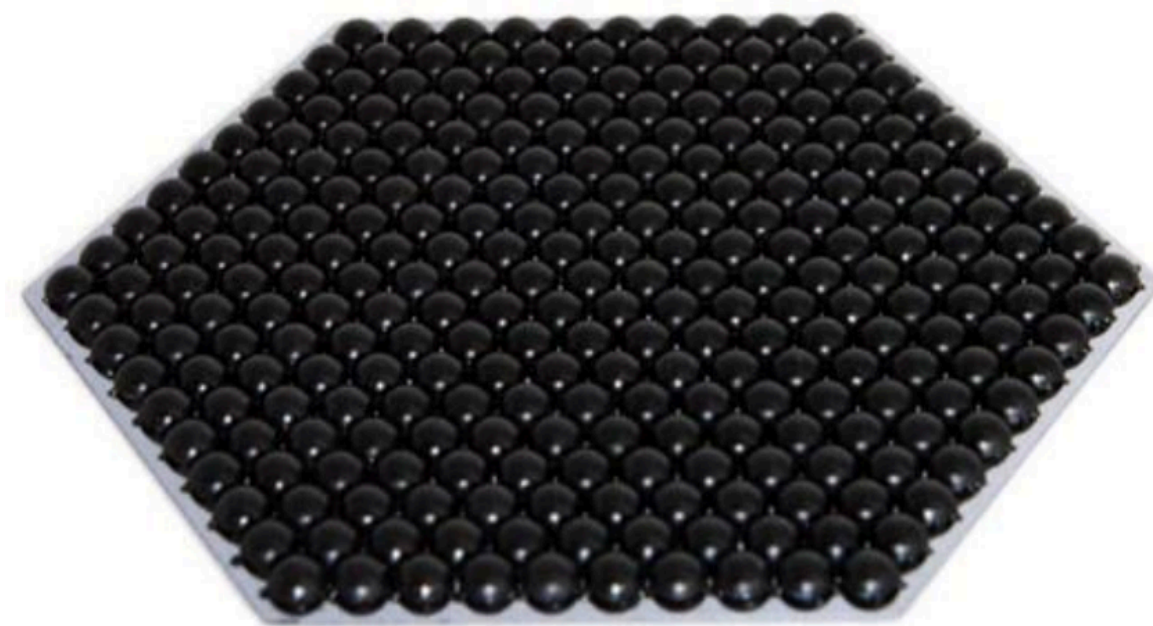


NIST: OMT technology,
feed-horn: ACT, SPT,
SO, LiteBIRD (High
frequency)

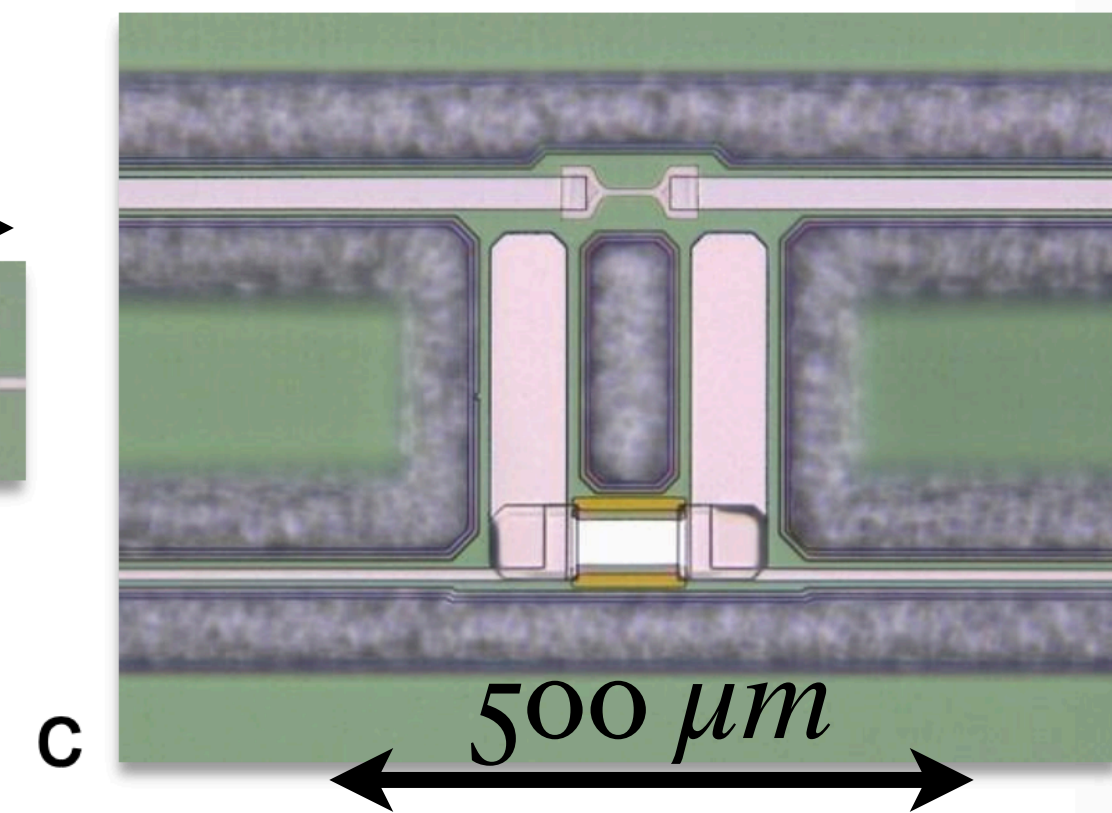
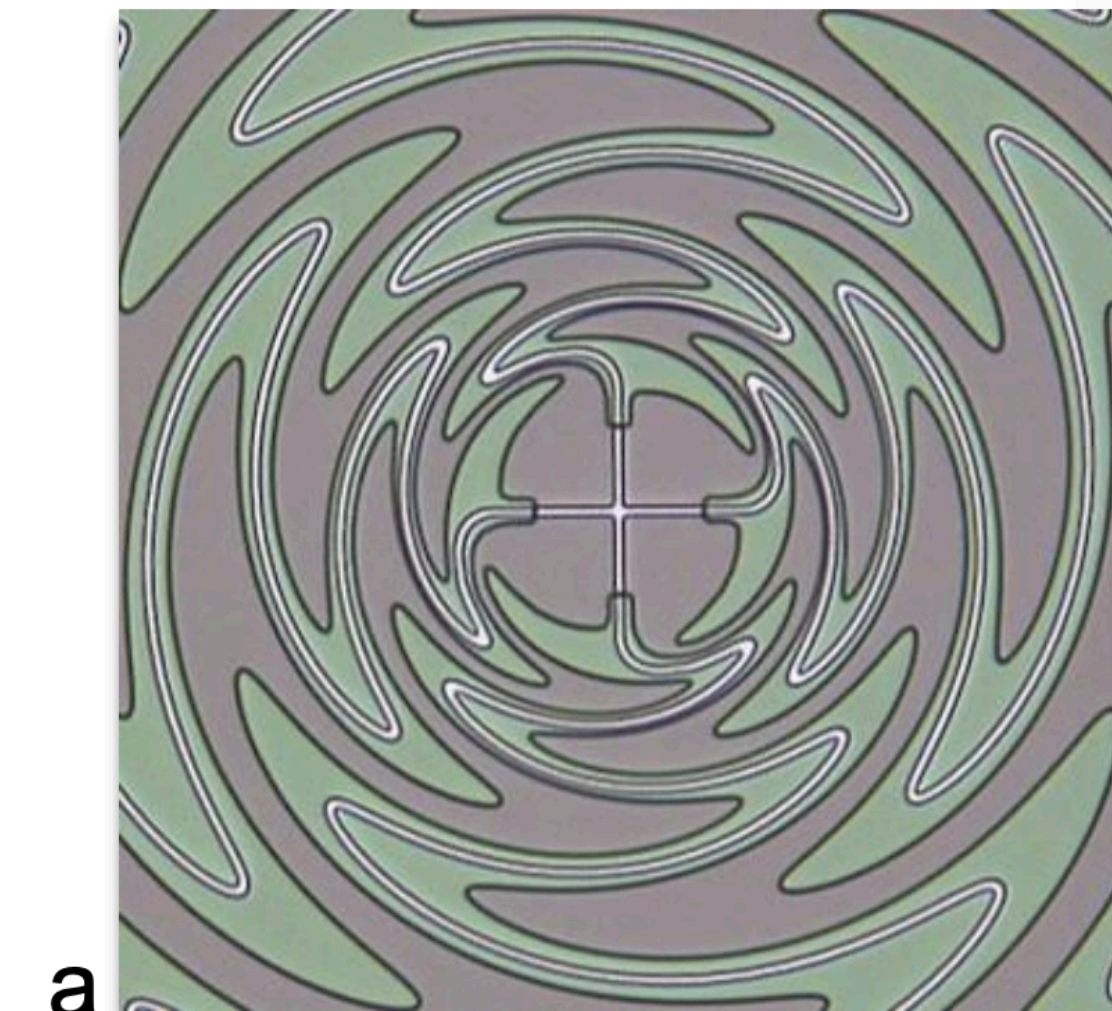
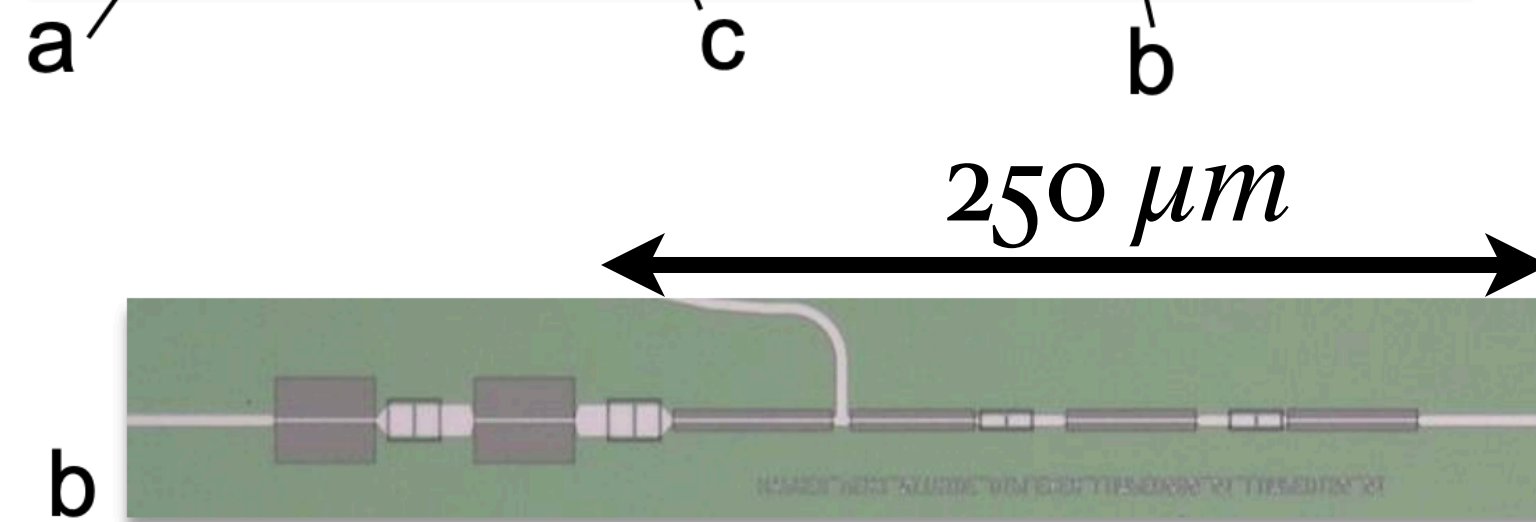
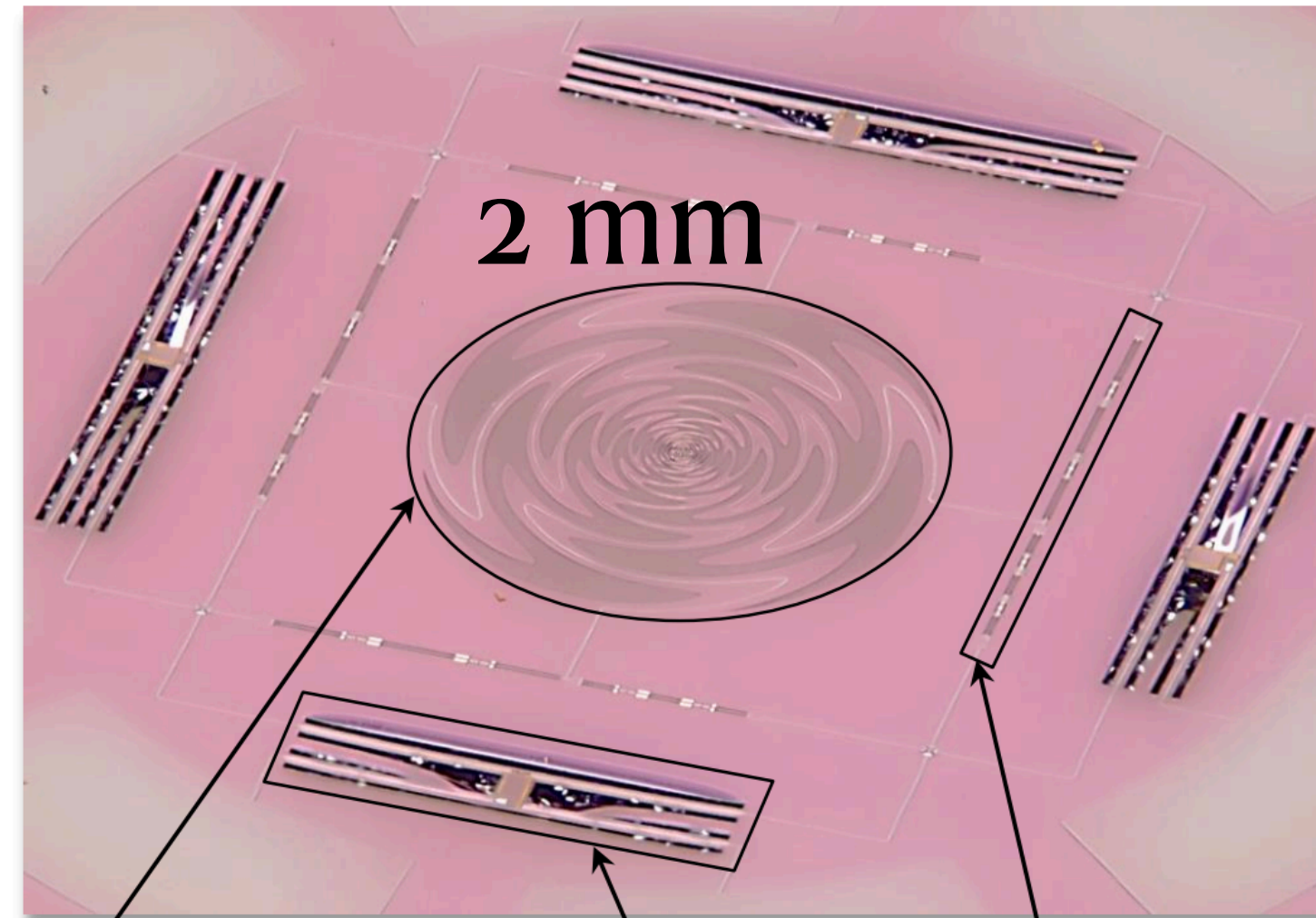
Lenslet-couple sinuous antenna TES



AR-coated silicon lenslet



Silicon lenslet array



- a) Sinuous antenna collects photons over wide frequency range.
- b) On-chip filters which separate into two frequency bands (e.g. 90/150 GHz).
- c) TES, superconducting AlMn (160 mK).

TES electronic system

- TES operates at very low temperature, electronic system to read data from thousand of TES is a challenge.
- There are a lot of developments of electronic system.
- Need careful tests at the laboratory.

Readout chain

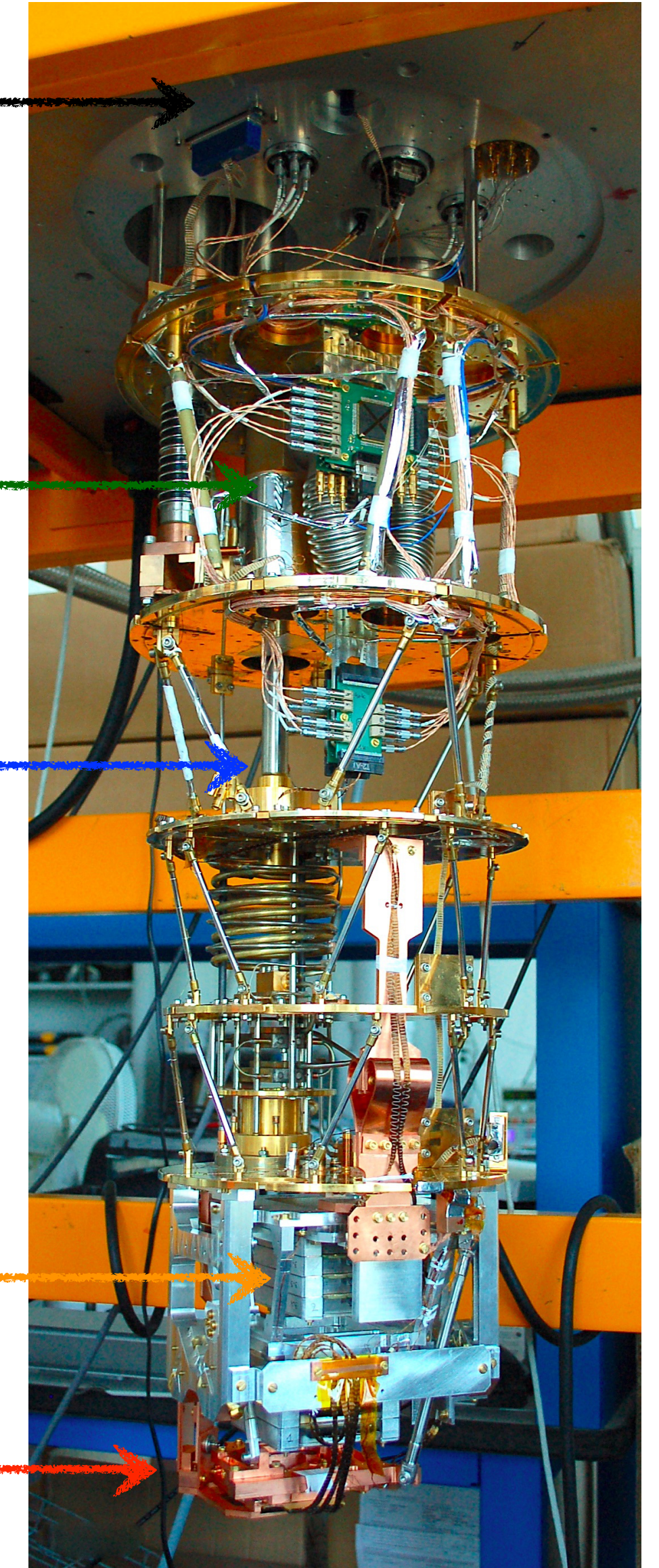
27°C~300 K

40 K

4 K

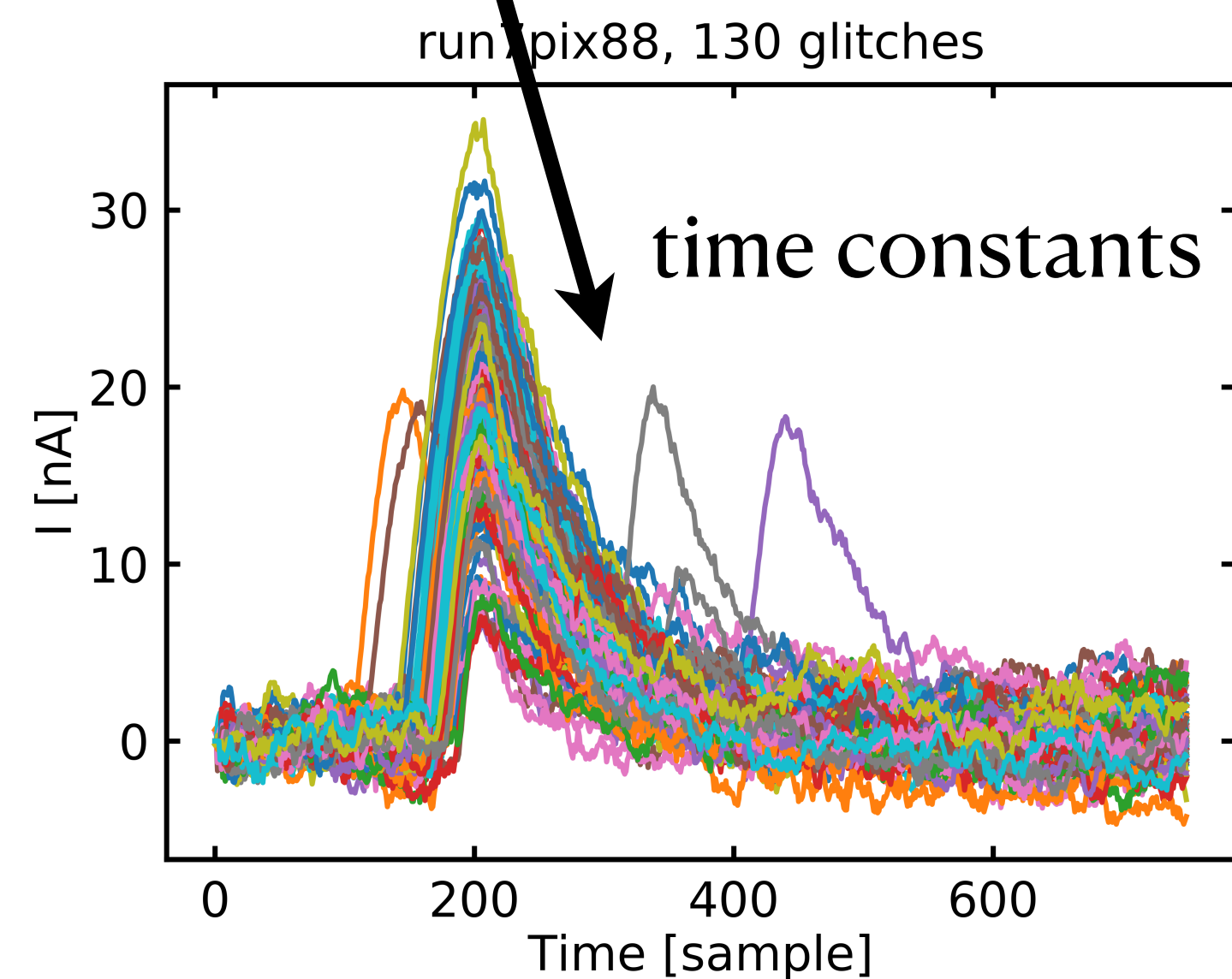
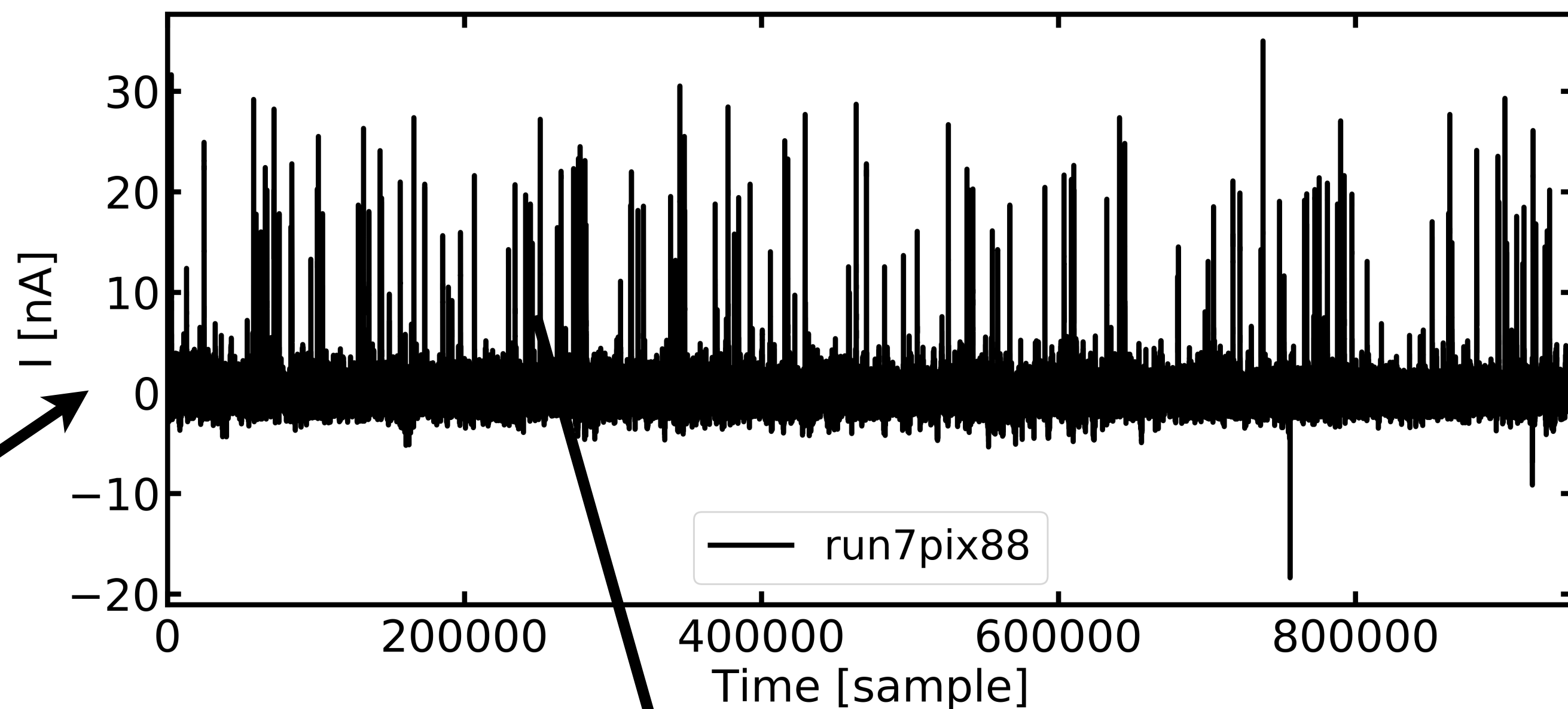
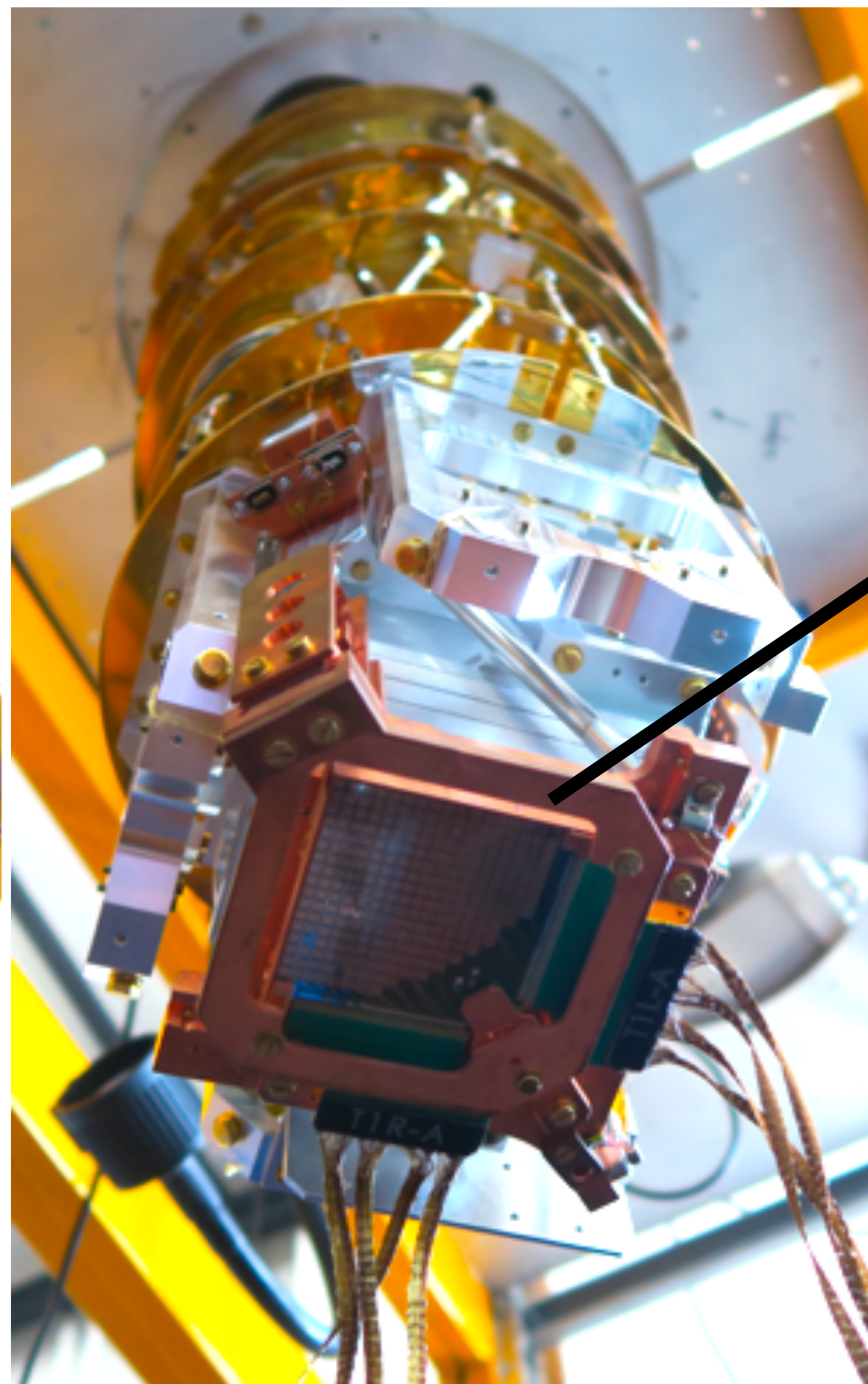
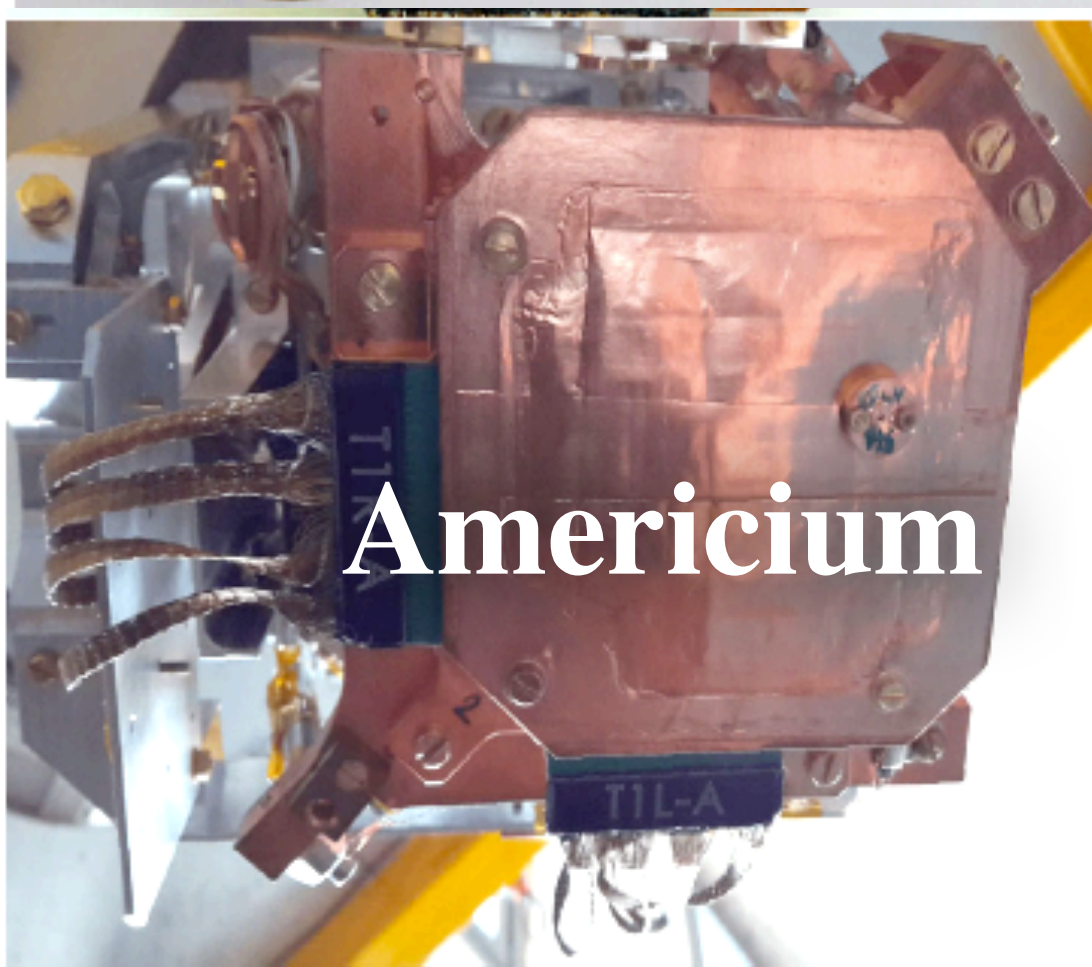
1 K

10 mK



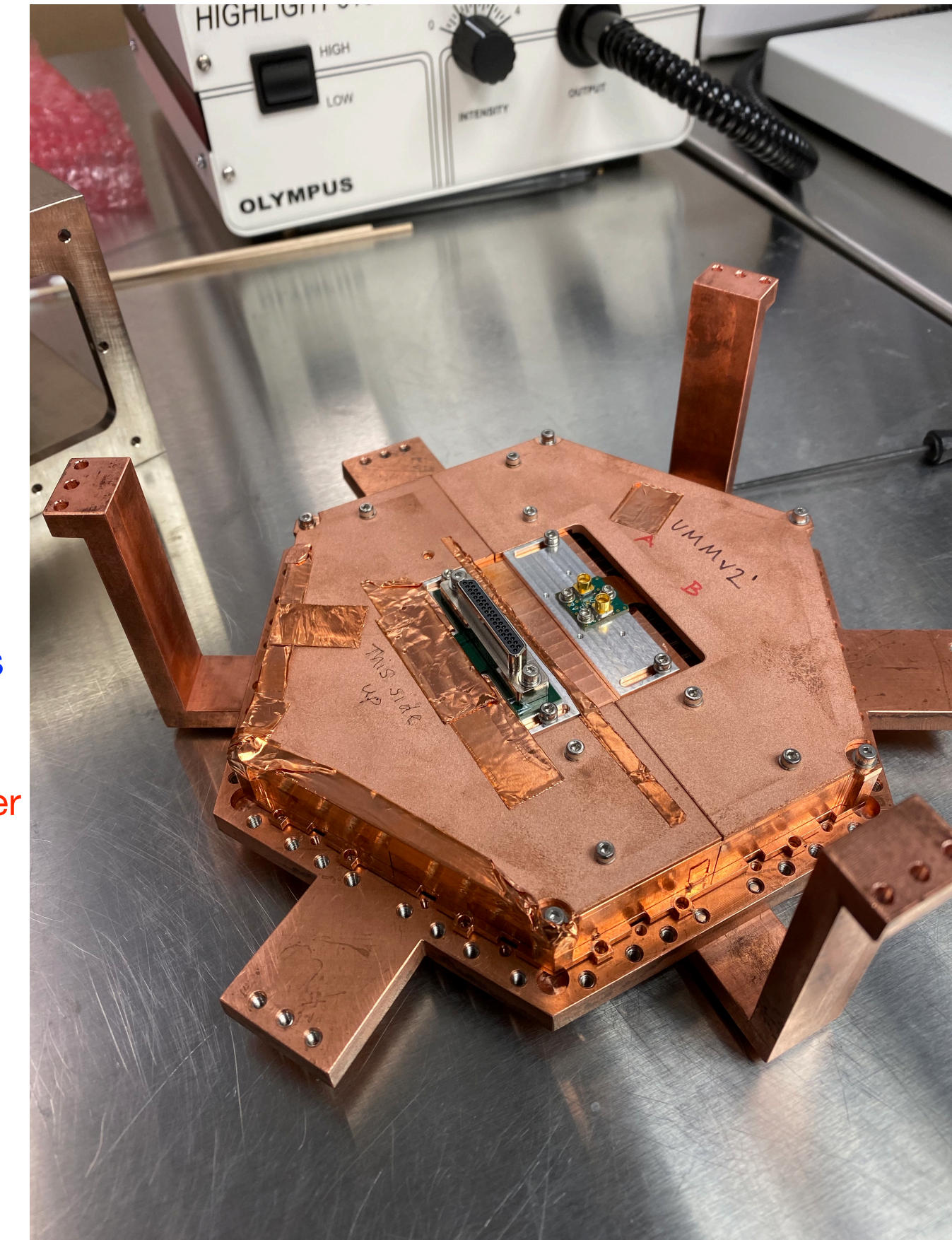
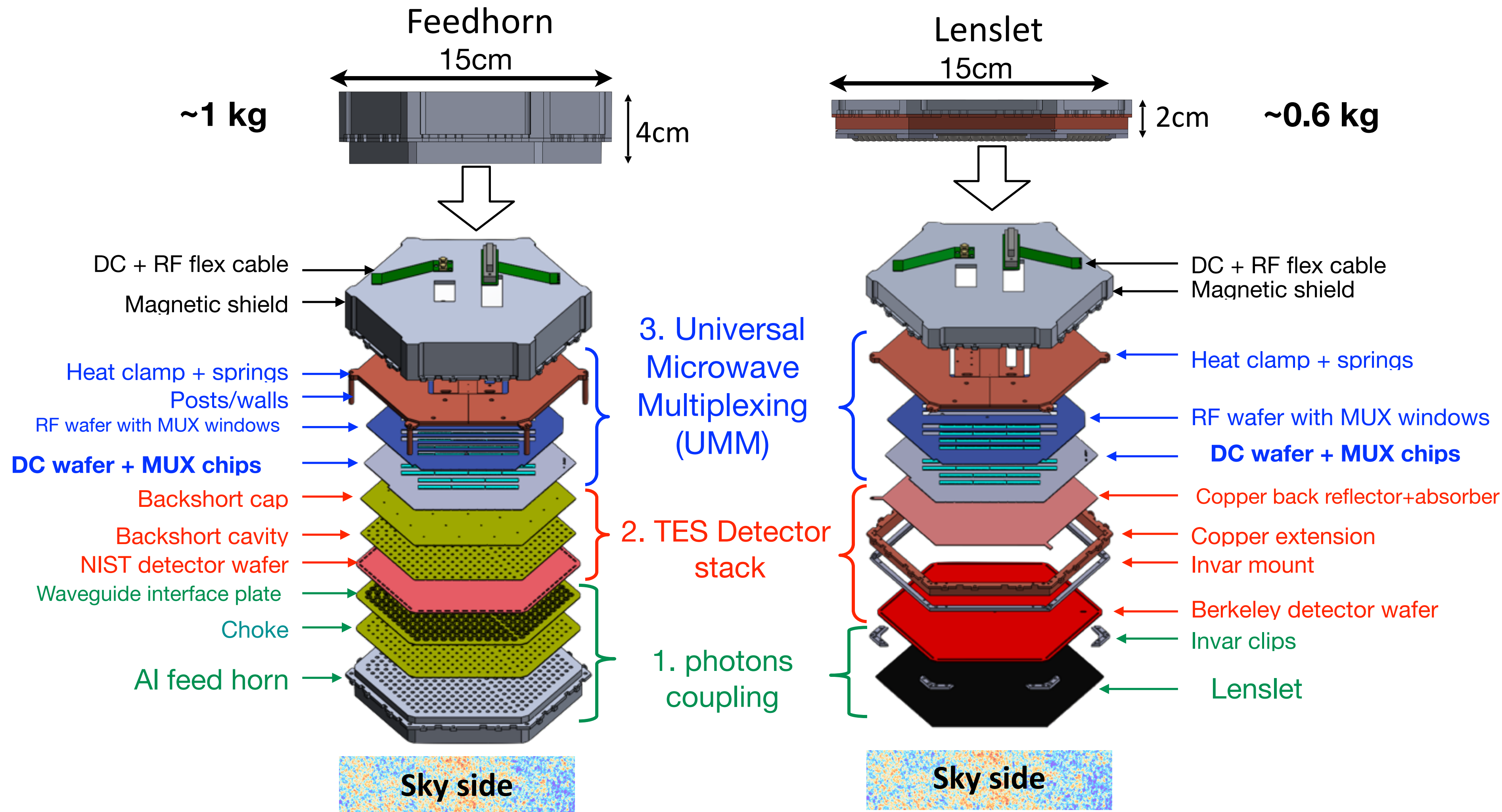
Cryogenics system

Testbed example on ground: Particles with Transition Edge Sensor array.

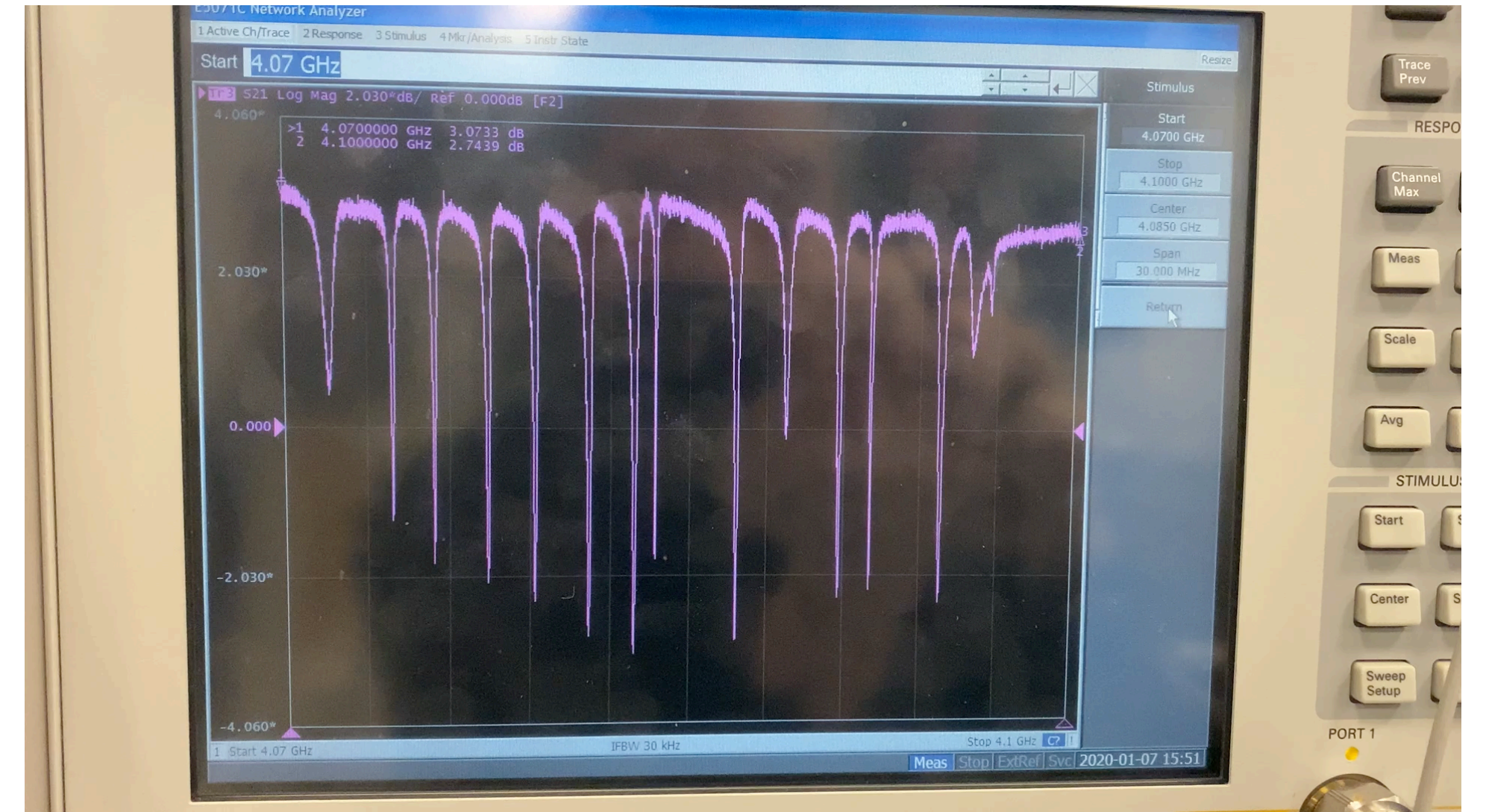
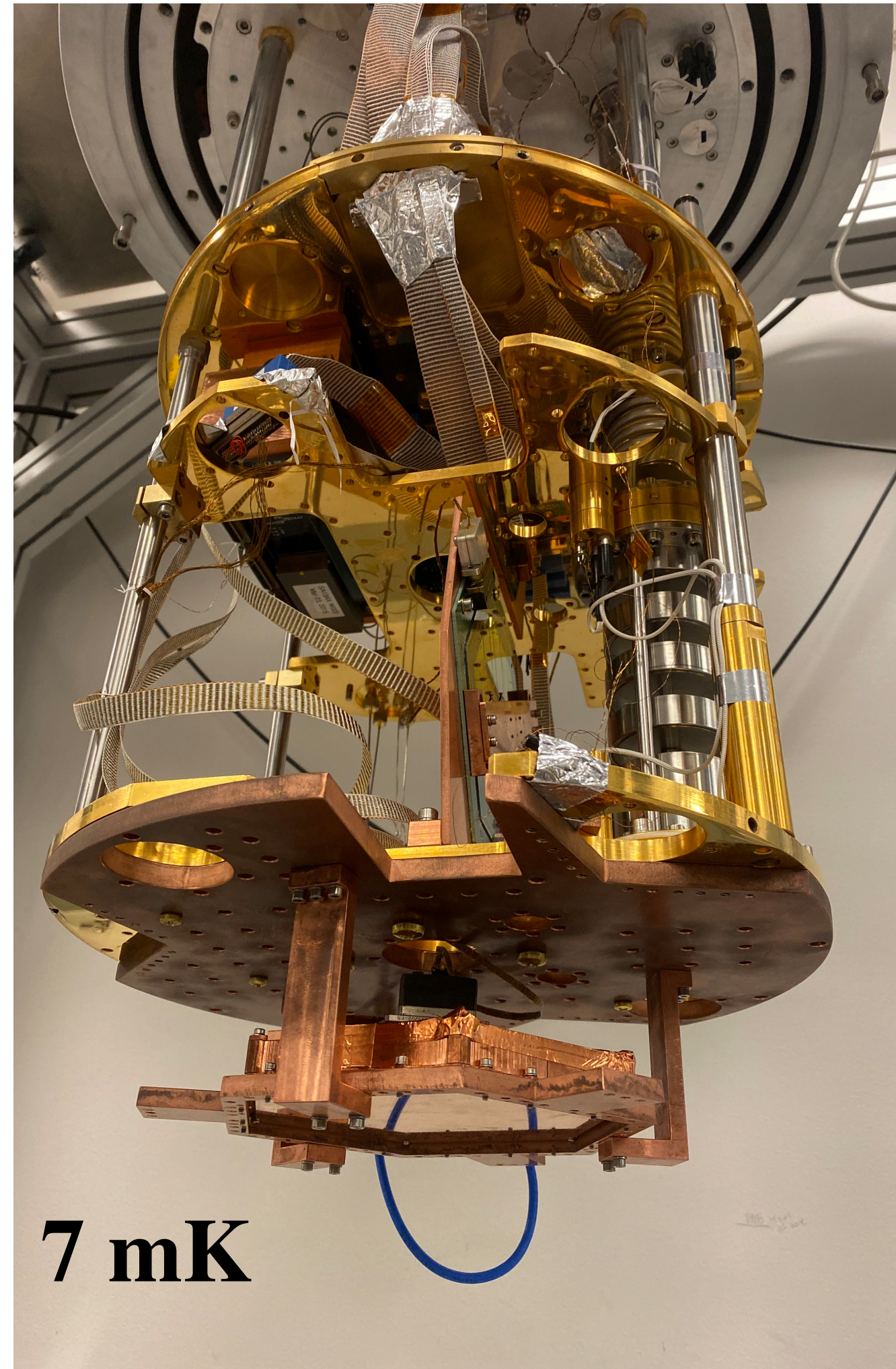
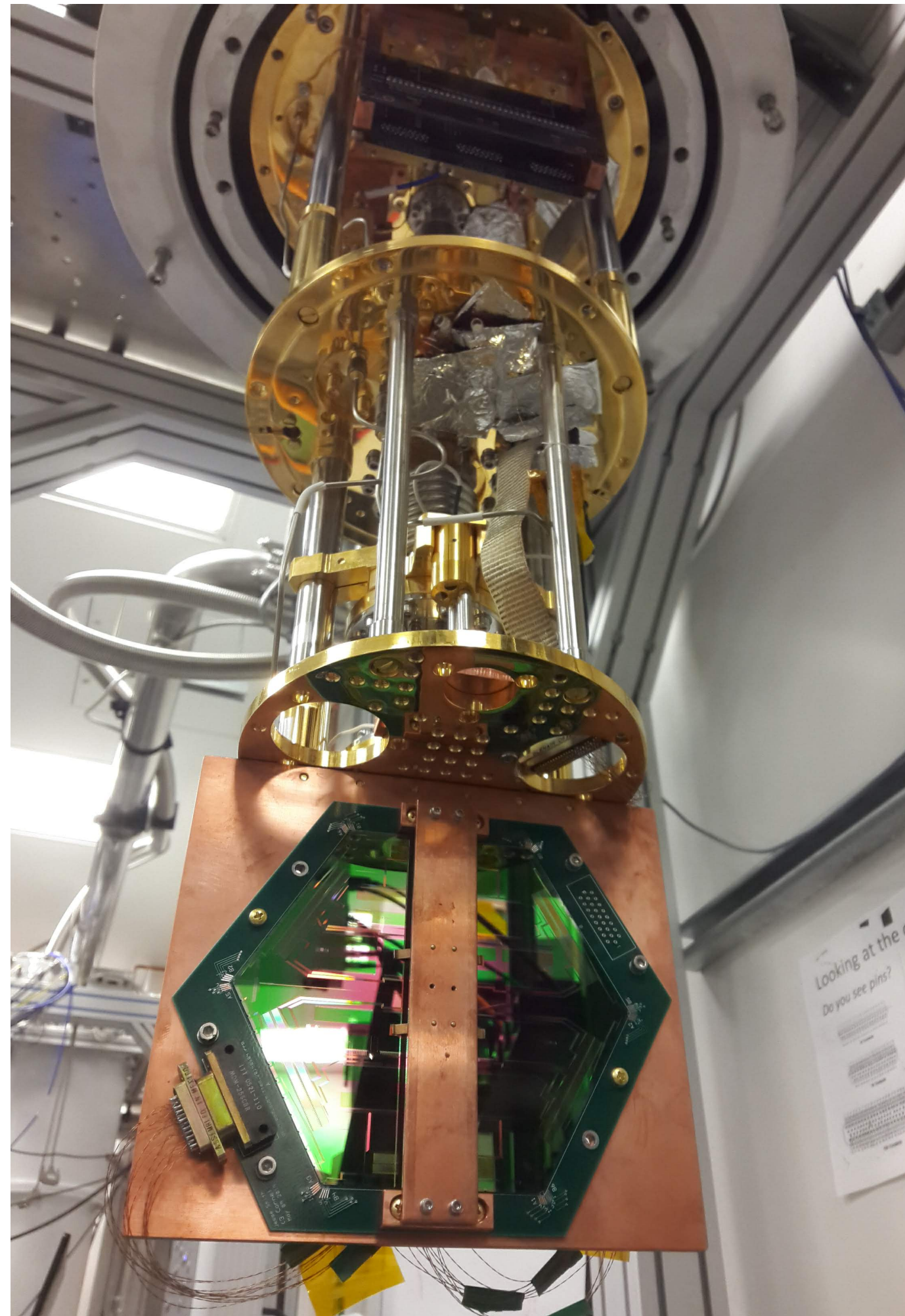


- 256 pixels array with radioactive Americium.
- When particles hit the TES, similarly CMB photons hit TES.

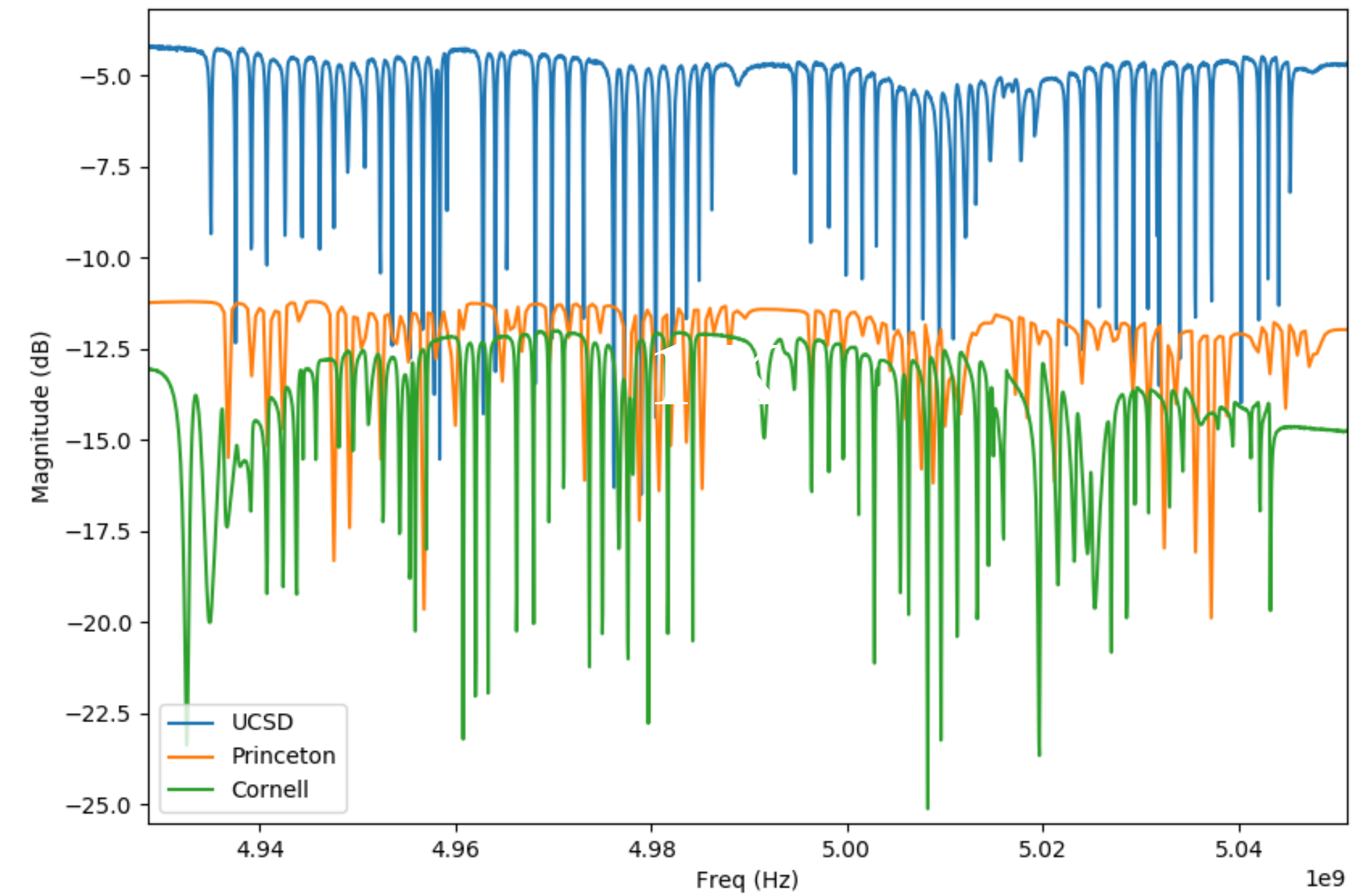
Testbed example on ground for TES array



- **Many stacked wafers to improve bolometer efficiency**
- The Transition Edge Sensor (TES) will be AlMn alloys with critical temperature 160 mK
- ~ 2000 TES per detector wafer, the normal resistance target is 8 mΩ

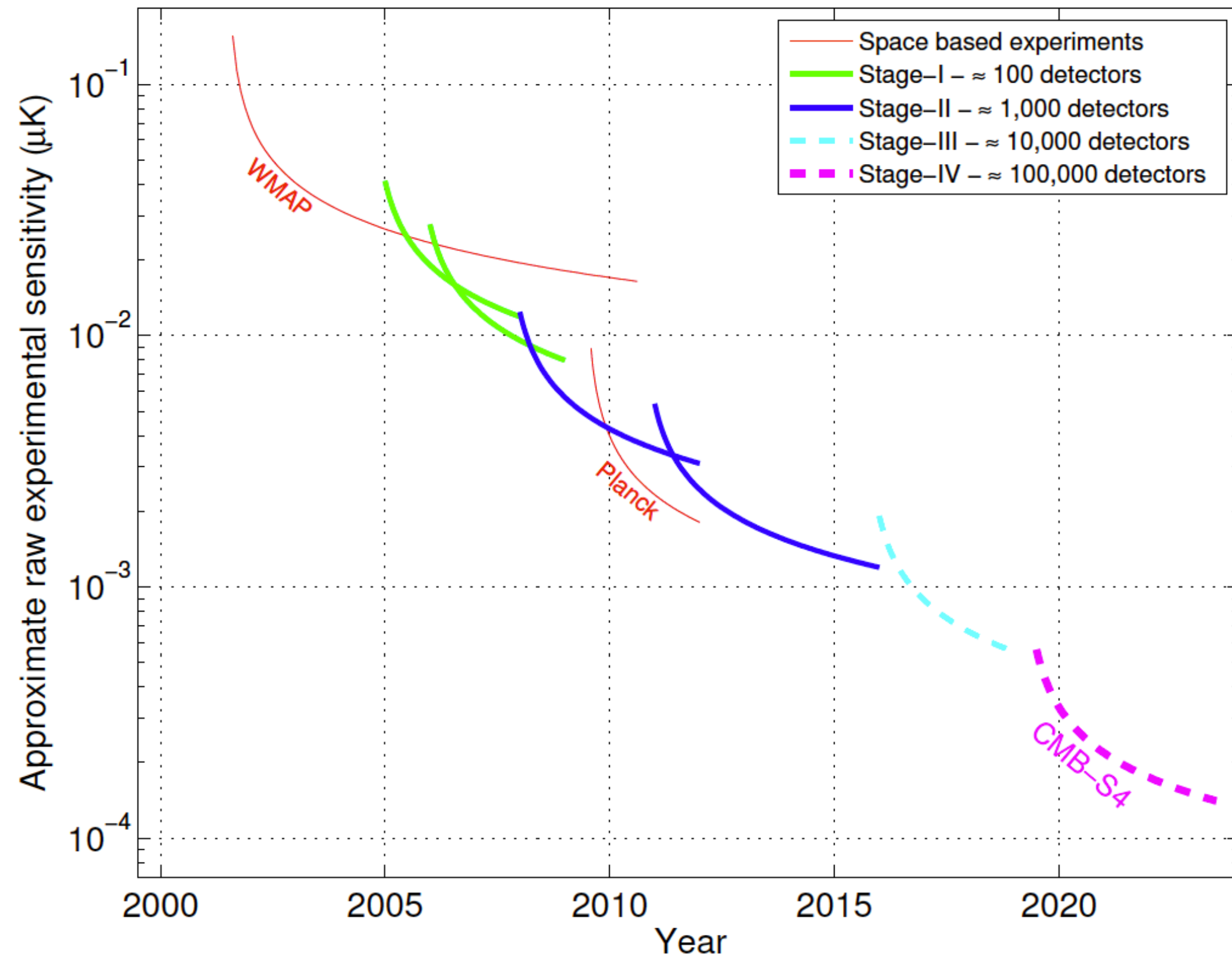


MMBv0a - A Retrospective



- Characterize the focal plane system at Cornell U.

Summary



- TES is a very sensitive detector operating at very low temperature.
- TES readout is a complex system that requires a lot of testing, characterizations, especially for a satellite mission.
- We don't have TES testbed facilities in Vietnam. CMB-Inflate project brings USTH-researchers a chance to work at advanced institutes in Europe, Japan.

- Detector sensitivity has been limited by photon shot noise. Increase number of detector to improve sensitivity.